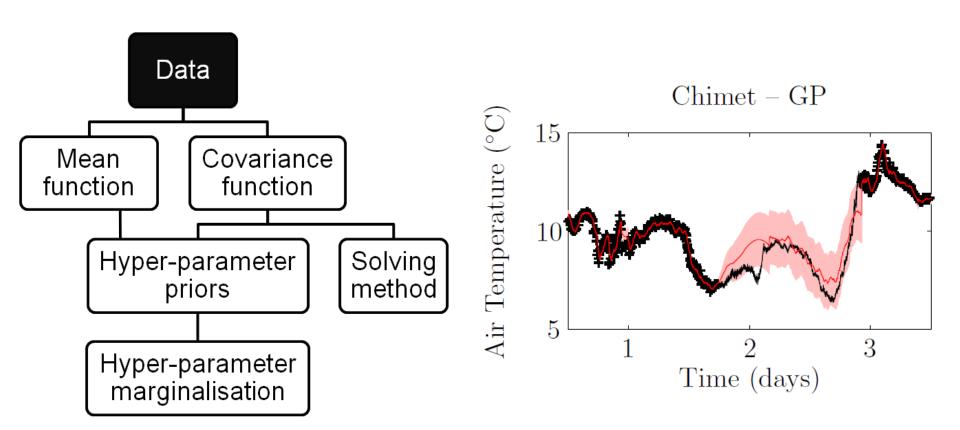
## An Introduction to Fitting Gaussian Processes to Data

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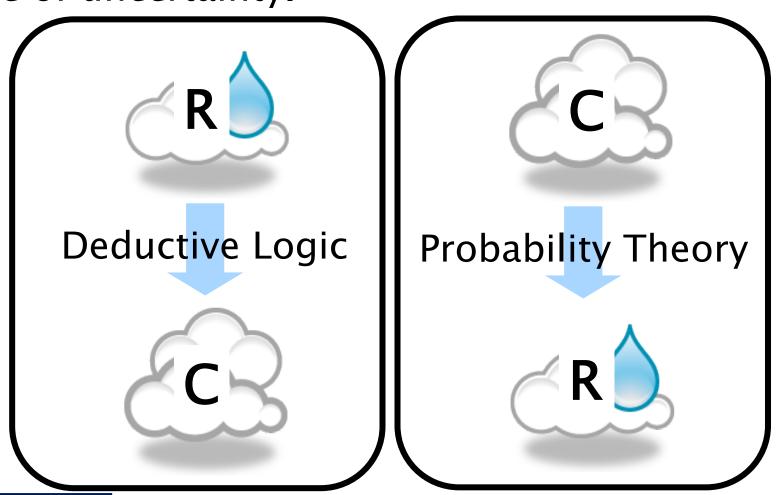


### You will learn how to fit a Gaussian process to data.



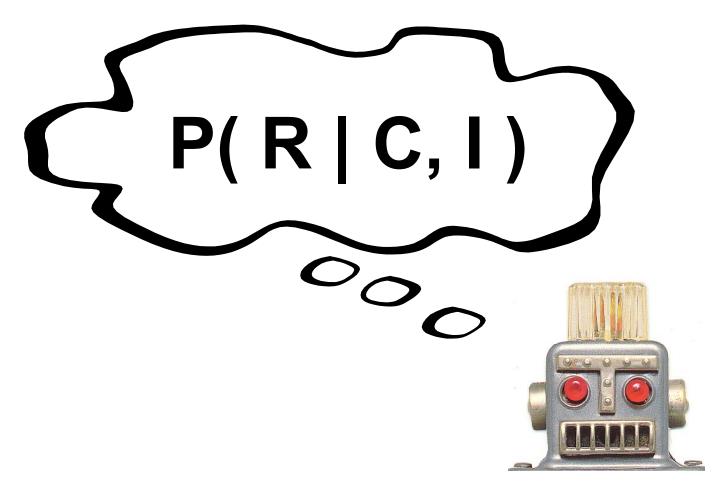


Probability theory represents an extension of traditional logic, allowing us to reason in the face of uncertainty.



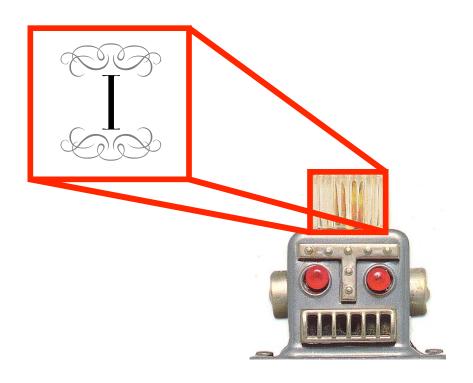


A probability is a degree of belief. This might be held by any agent - a human, a robot, a sensor, etc.





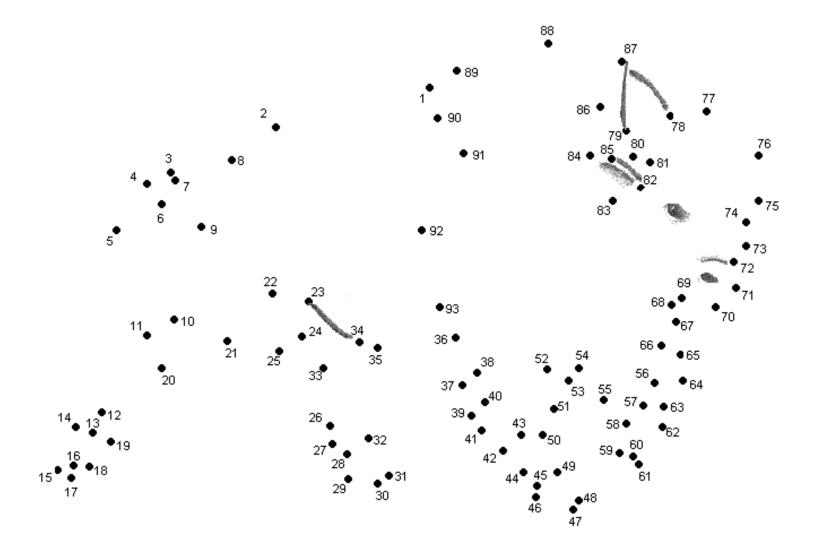
'I' is the totality of an agent's prior information. An agent is defined by *I*.



We define our agents so that they can perform difficult inference for us.

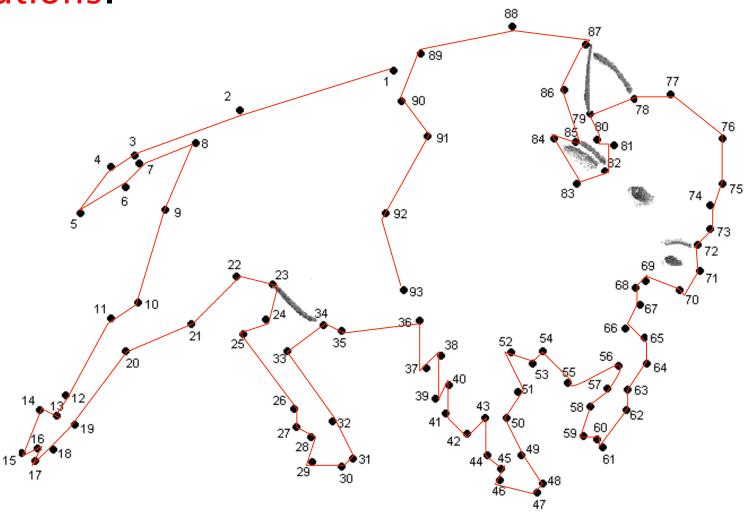


#### A dot-to-dot is an inference problem.



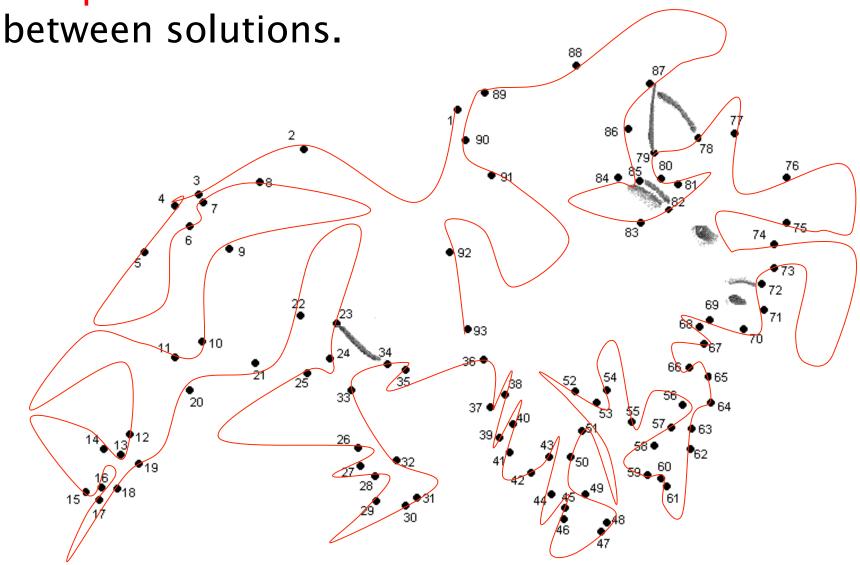


## A dot-to-dot is a problem with many possible solutions.



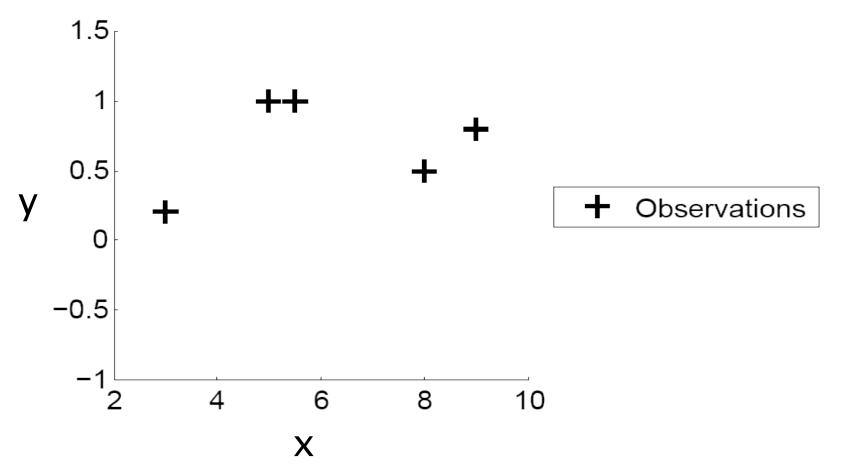


Our prior information allows us to discriminate





A dot-to-dot requires us to do inference about functions, as can be seen more clearly in one dimension.



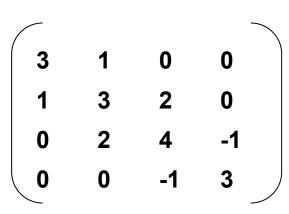


The multivariate Gaussian distribution is wonderful; it is defined by a mean vector (which simply gives the centrepoint) and covariance matrix.

mean  $\mathbf{y}_1$ 



The covariance K must be a positive semidefinite matrix; so for any vector x,  $x^T$  K  $x \ge 0$ . This implies that:



*K* must be symmetric.

The diagonal of *K* must be positive.

 $K = R^T R$  for some upper triangular R.

The eigenvalues of K are all  $\geq 0$ .



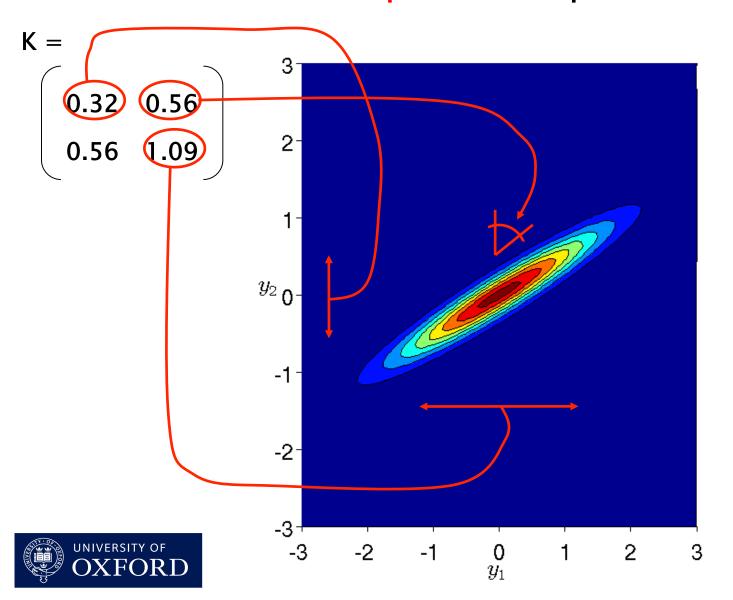
We can represent any covariance *K* using the spherical parameterisation.

$$K = R^{T}R$$

$$R = \begin{pmatrix} 1 & \cos(\theta_{1}) & \cos(\theta_{2}) & \cdots \\ 0 & \sin(\theta_{1}) & \sin(\theta_{2})\cos(\theta_{3}) \\ 0 & 0 & \sin(\theta_{2})\sin(\theta_{3}) \\ \vdots & & \ddots \end{pmatrix} \begin{pmatrix} h_{1} & 0 & 0 & \cdots \\ 0 & h_{2} & 0 \\ 0 & 0 & h_{3} \\ \vdots & & \ddots \end{pmatrix}$$



The (i,j)<sup>th</sup> element of the covariance expresses how variable i is dependent upon variable j.



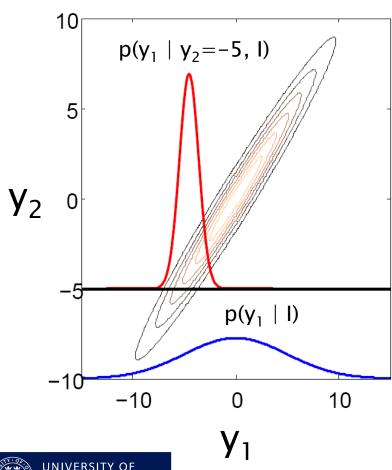
The (i,j)<sup>th</sup> element of the inverse covariance (precision) expresses how variable i is dependent upon variable j, conditioned on all other variables.

$$K = \begin{bmatrix} 1.67 & -2 & 1.33 \\ -2 & 3 & -2 \\ 1.33 & -2 & 1.67 \end{bmatrix}$$

$$\mathbf{K}^{-1} = \begin{bmatrix} 3 & 2 & 0 \\ 2 & 3 & 2 \\ 0 & 2 & 3 \end{bmatrix} \quad 0 \qquad y_2 \qquad y_3$$

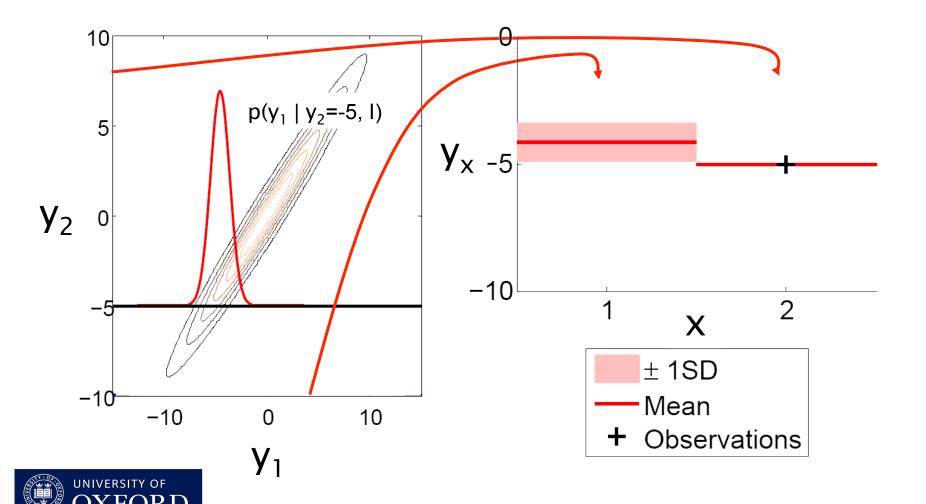


# The multivariate Gaussian distribution allows us to produce distributions for variables conditioned on any other observed variables.

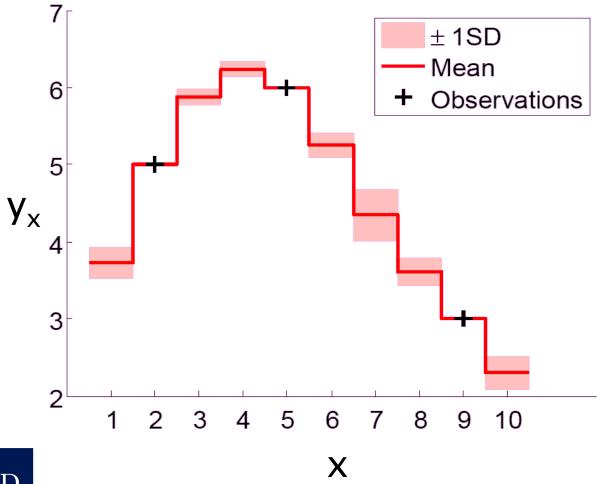




The Gaussian distribution allows us to produce distributions for variables conditioned on any other observed variables.



A Gaussian process is the generalisation of a multivariate Gaussian distribution to a potentially infinite number of variables.





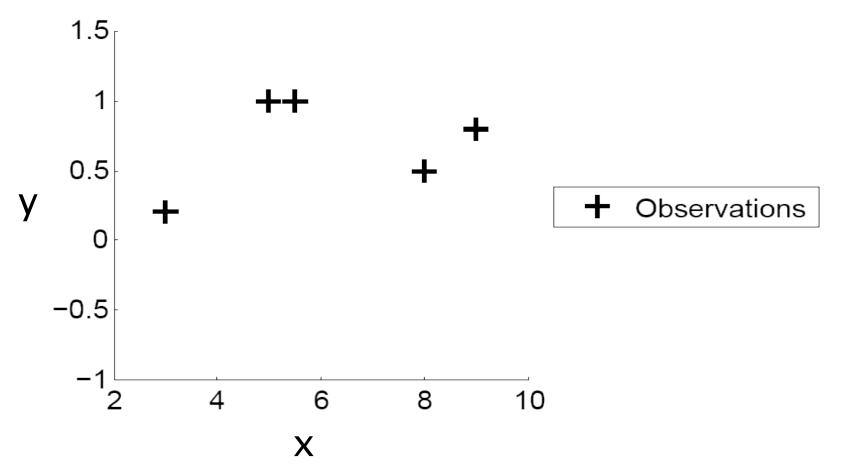
For a Gaussian process, we need to define mean and covariance functions, specified by hyperparameters  $\phi$ .

$$\begin{pmatrix} \mu_1 \\ \mu_2 \\ \mu_3 \\ \vdots \end{pmatrix} = \begin{pmatrix} \mu(x_1; \phi) \\ \mu(x_2; \phi) \\ \mu(x_3; \phi) \\ \vdots \end{pmatrix}$$

$$\begin{pmatrix} K_{11} & K_{21} & K_{13} & \cdots \\ K_{21} & K_{22} & K_{23} \\ K_{31} & K_{32} & K_{33} \\ \vdots & & \ddots \end{pmatrix} = \begin{pmatrix} K(x_1, x_1; \phi) & K(x_1, x_2; \phi) & K(x_1, x_3; \phi) & \cdots \\ K(x_2, x_1; \phi) & K(x_2, x_2; \phi) & K(x_2, x_3; \phi) \\ K(x_3, x_1; \phi) & K(x_3, x_2; \phi) & K(x_3, x_3; \phi) \\ \vdots & & \ddots \end{pmatrix}$$

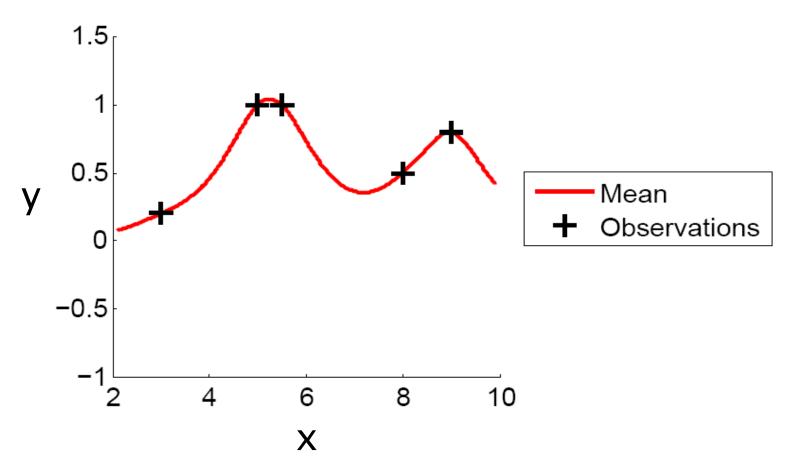


### A Gaussian process represents a powerful way to perform Bayesian inference about functions.



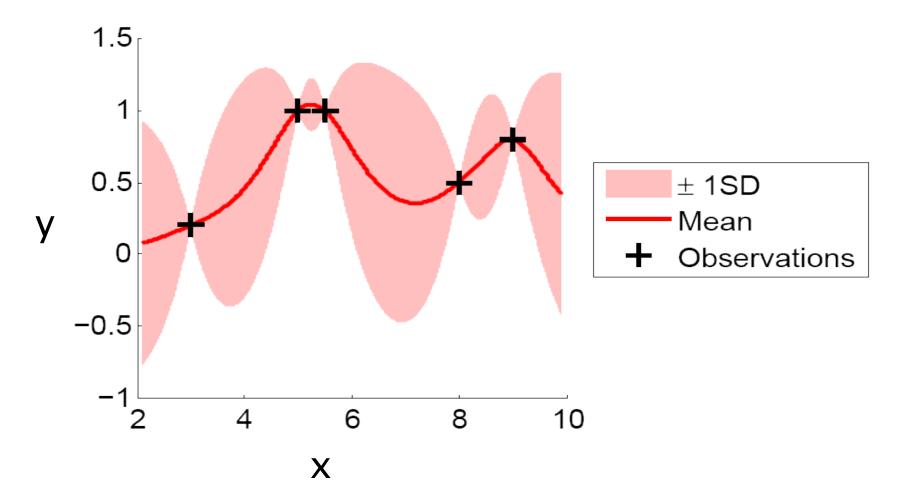


### A Gaussian process produces a mean estimate.





## A Gaussian process produces a mean estimate along with an indication of the uncertainty in it.





### The posterior mean and covariance equations follow simply from Gaussian identities.

$$y_* = y(x_*)$$
 Predictants  
 $y_d = y(x_d)$  Data

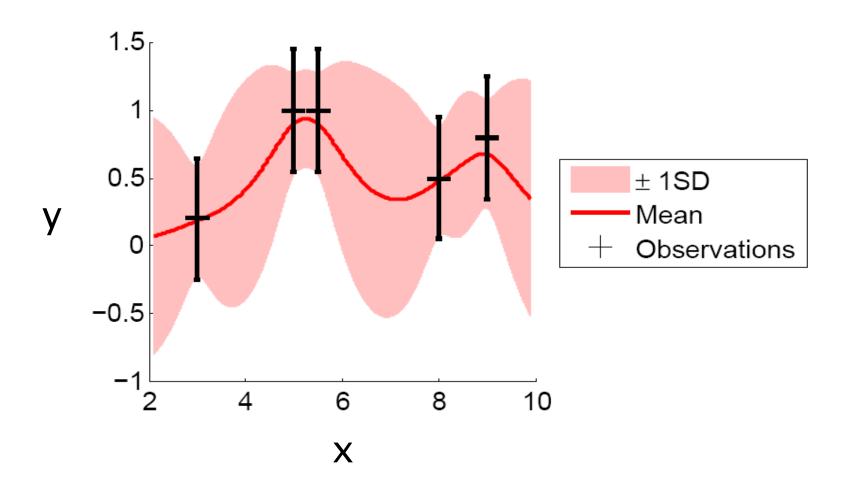
$$p(y_*, y_d) = N\left( \begin{pmatrix} y_* \\ y_d \end{pmatrix}; \begin{pmatrix} \mu(x_*) \\ \mu(x_d) \end{pmatrix}, \begin{pmatrix} K(x_*, x_*) & K(x_*, x_d) \\ K(x_d, x_*) & K(x_d, x_d) \end{pmatrix} \right)$$

Mean 
$$m(y_* | y_d) = \mu(x_*) + K(x_*, x_d) K(x_d, x_d)^{-1} (y_d - \mu(x_d))$$
  
Cov.  $C(y_* | y_d) = K(x_*, x_*) + K(x_*, x_d) K(x_d, x_d)^{-1} K(x_d, x_*)$ 



All functions here are dependent upon hyperparameters.

#### A Gaussian process can accommodate noise.





We usually consider making independent and identically distributed (IID) Gaussian noisy measurements z, of y; giving

$$p(z_d | y_d) = N(z_d; y_d, \sigma^2 I_d)$$
 Identity matrix

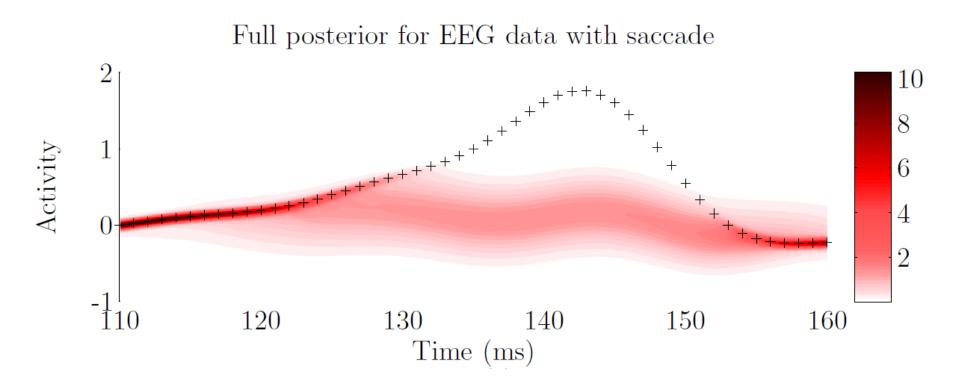
$$m(y_* | z_d) = \mu(x_*) + K(x_*, x_d) V(x_d, x_d)^{-1} (z_d - \mu(x_d))$$

$$C(y_* | z_d) = K(x_*, x_*) + K(x_*, x_d) V(x_d, x_d)^{-1} K(x_d, x_*)$$

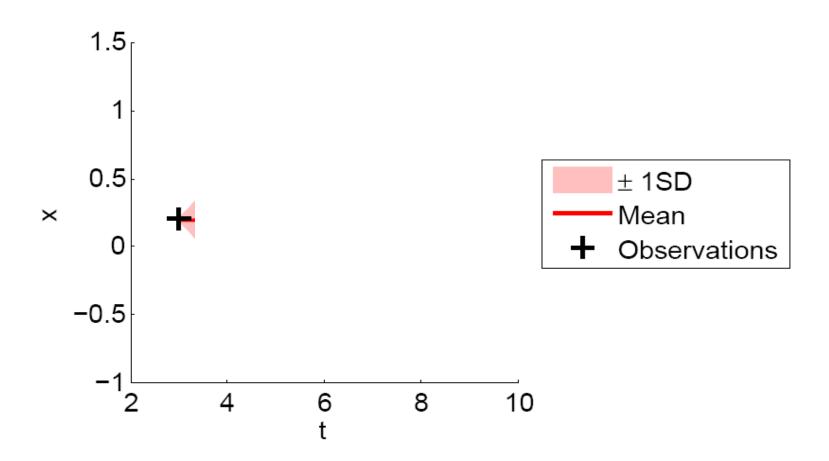
$$V(x_d, x_d) = K(x_d, x_d) + \sigma^2 I_d = \begin{pmatrix} K(x_1, x_1) + \sigma^2 & K(x_1, x_2) & \cdots \\ K(x_2, x_1) & K(x_2, x_2) + \sigma^2 \\ \vdots & \ddots \end{pmatrix}$$



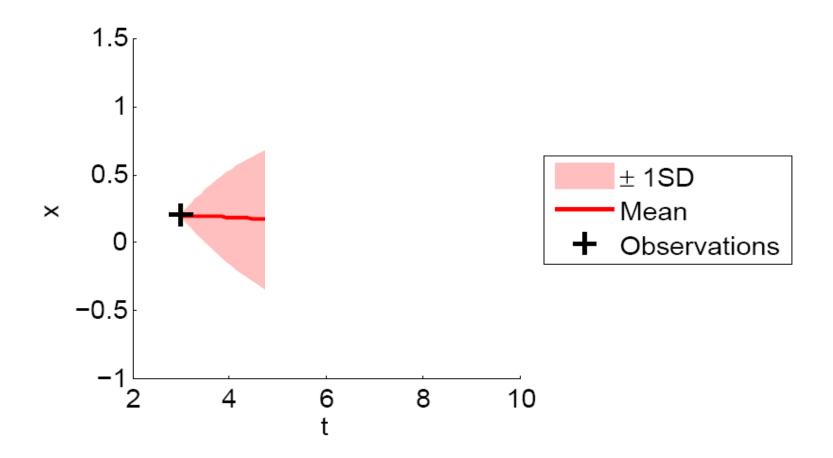
## More generally, we could consider correlated noise, in which the noise contribution could itself be drawn from a GP.



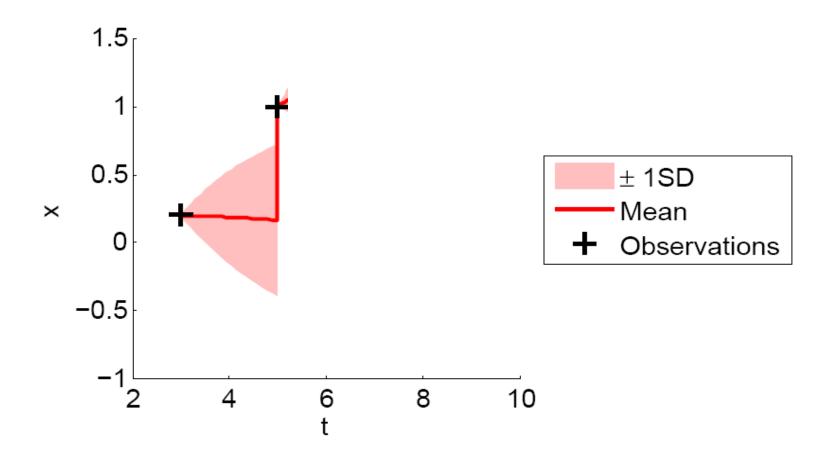




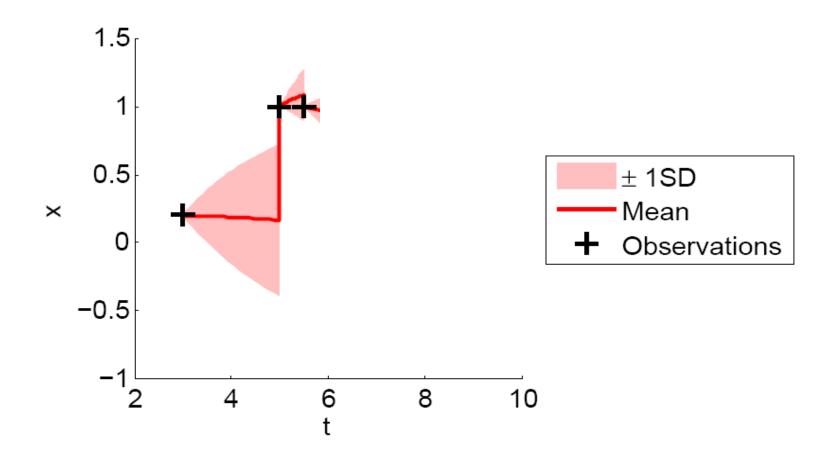




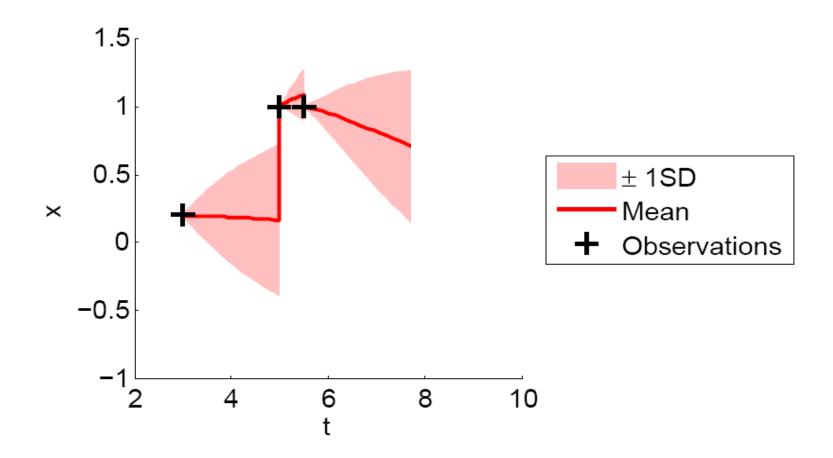




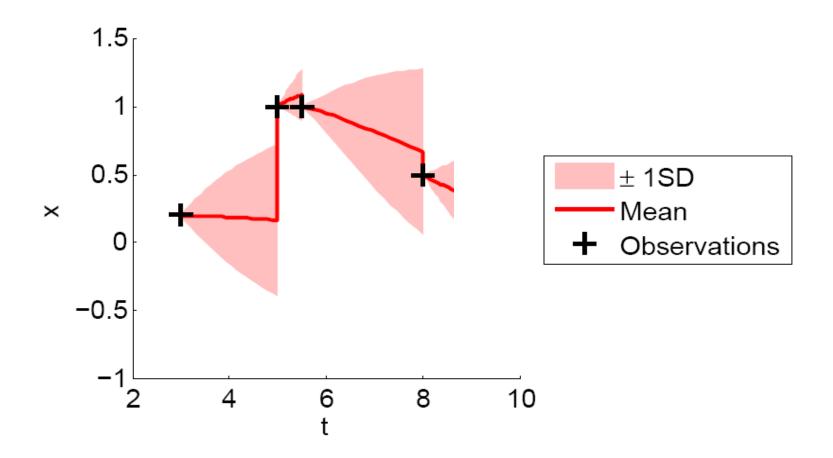




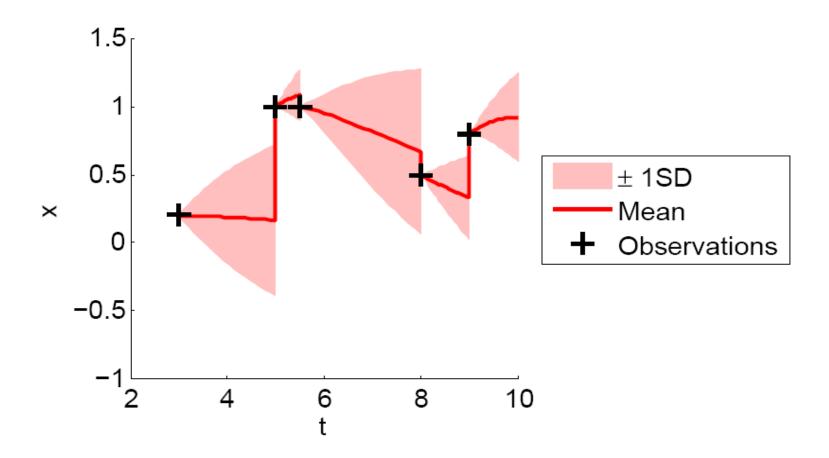




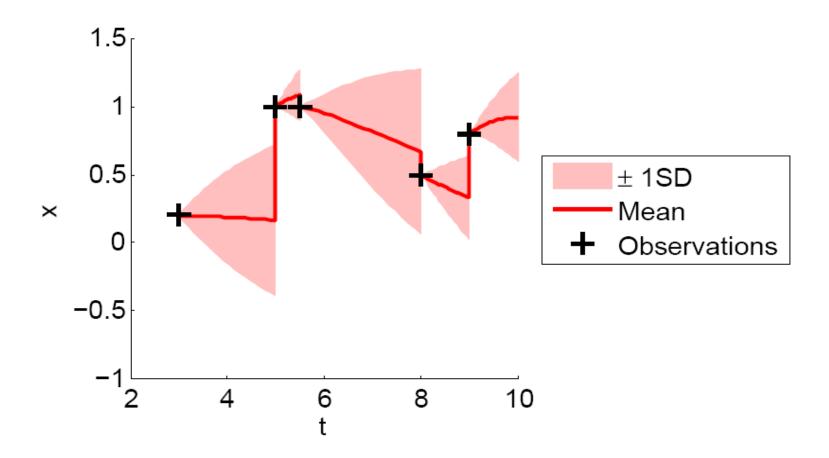






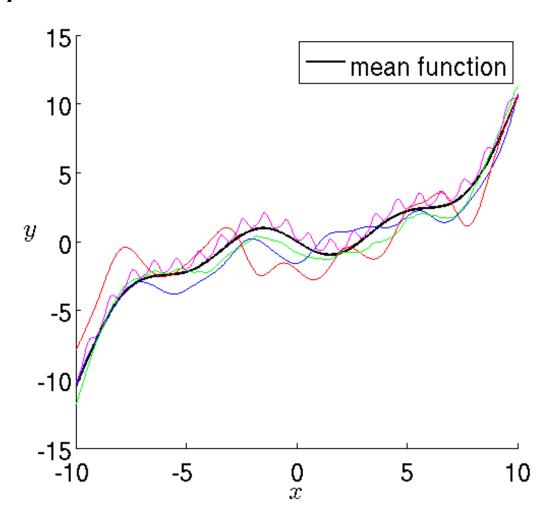






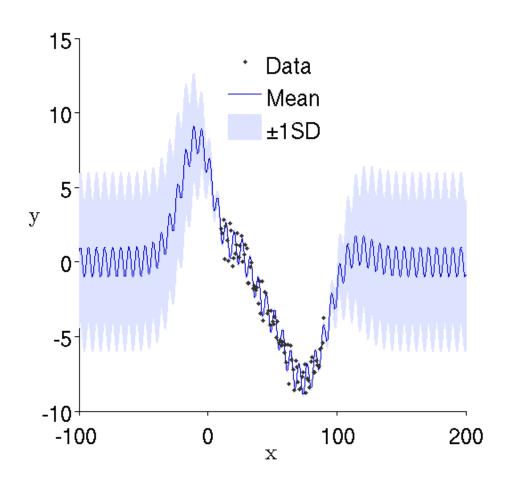


The prior mean function  $\mu(x;\phi)$  should be our best guess (of any form) for the function y(x) before any observations are made.





### The prior mean function is the function our inference will default to far from observations.





It's rarely worth using a complicated mean function (with many hyperparameters), unless we're concerned with prediction far from our observations.

Predictions required	Mean function
Interpolation	$\mu(x;\phi) = \text{mean}(y_d)$ .
Extrapolation	Bespoke model built using domain knowledge.



There are a huge number of covariance functions (in spite of the requirement that they be positive semi-definite) appropriate for modelling functions of different types.

Function type	Covariance function
Improbably smooth	Squared exponential.
Less smooth	Matérn.
Polynomial	Polynomial.



Many covariance functions (including the squared exponential and Matérn) are of the metric form

squared output scale

$$K(x_i, x_j; w) = h^2 \kappa (d(x_i, x_j; w))$$

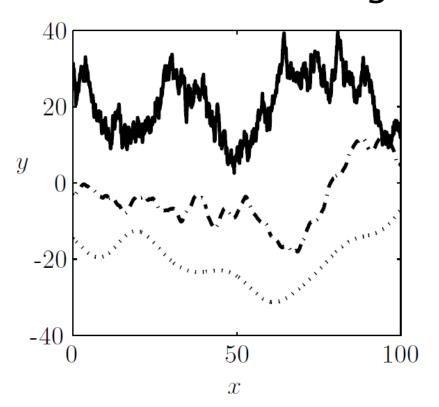
increasing d

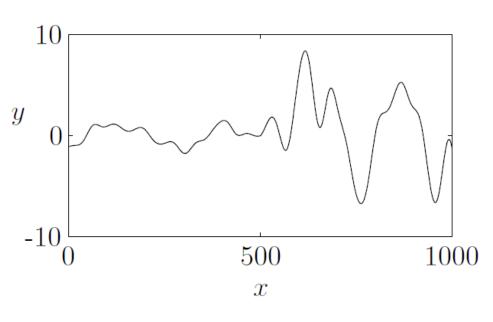
decreases with distance input function scale

e.g. 
$$d(x_i, x_j; w) = \frac{|x_i - x_j|}{w}$$



We often want distances that are stationary (a function of  $x_1$ – $x_2$ ), implying that the function looks similar throughout its domain.



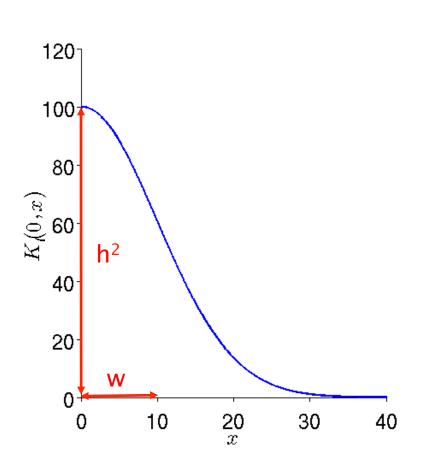


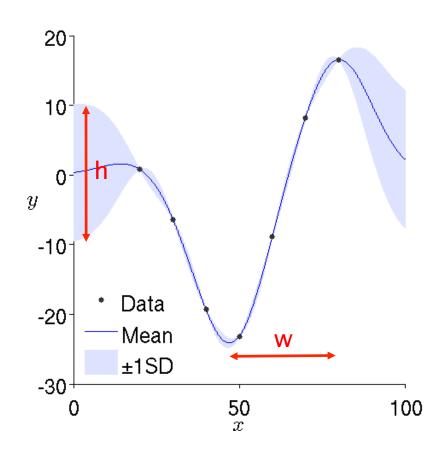
stationary functions

non-stationary function



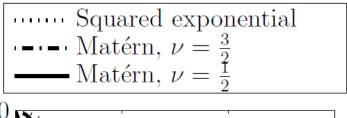
The hyperparameters *h* and *w* specify our expected length scales of the function in output ('height') and input ('width') spaces respectively.

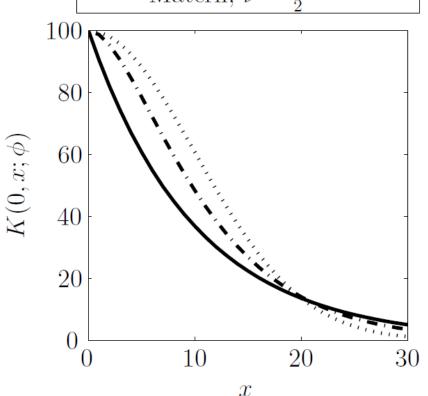


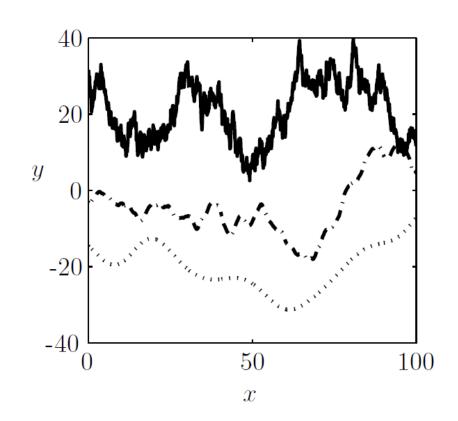




# The squared exponential and Matérn covariances allow us to model functions of various degrees of smoothness.







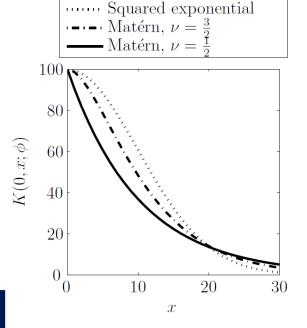


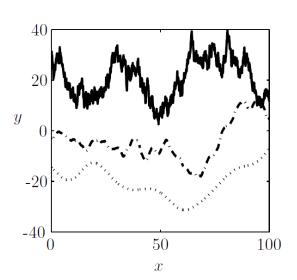
# The squared exponential and Matérn covariances allow us to model functions of various degrees of smoothness.

$$K_{\text{SE}}(x_{i}, x_{j}; h, w) = h^{2} \exp(-\frac{1}{2} d(x_{i}, x_{j}; w)^{2})$$

$$K_{\text{Mtn}}(x_{i}, x_{j}; h, w, v = \frac{3}{2}) = h^{2} (1 + \sqrt{3} d(x_{i}, x_{j}; w)) \exp(-\sqrt{3} d(x_{i}, x_{j}; w))$$

$$K_{\text{Mtn}}(x_{i}, x_{j}; h, w, v = \frac{1}{2}) = h^{2} \exp(-d(x_{i}, x_{j}; w))$$

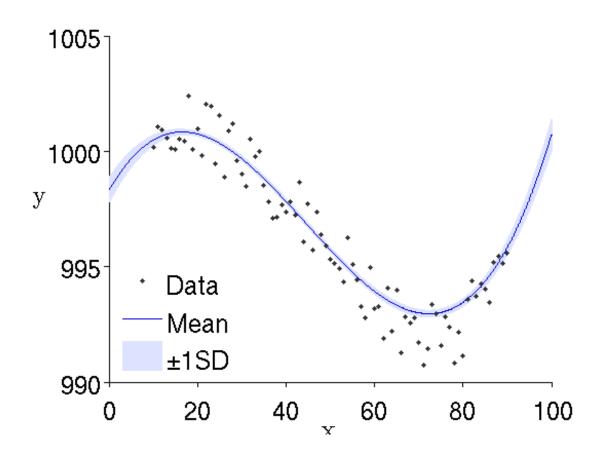






#### Polynomial covariances exist to model functions that are known to be polynomial.

$$K_{\mathrm{P}}(x_i, x_j; x, W) = (c^2 + x_i^T W x_j)^d$$



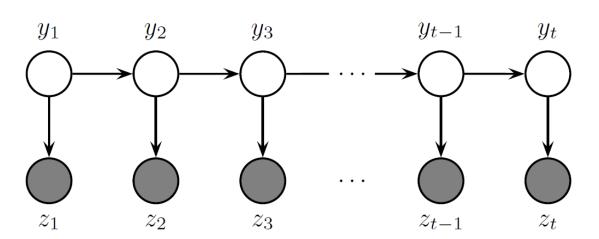


The Kalman filter is a Gaussian process with a

special covariance function, one that gives a sparse precision matrix. This allows efficient computation.

$$K = (ugly)$$

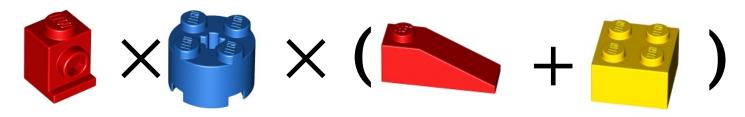
$$K^{-1} = \begin{pmatrix} 2 & -1 & 0 & 0 & \cdots \\ -1 & 2 & -1 & 0 \\ 0 & -1 & 2 & -1 \\ 0 & 0 & -1 & 2 \\ \vdots & & \ddots \end{pmatrix}$$





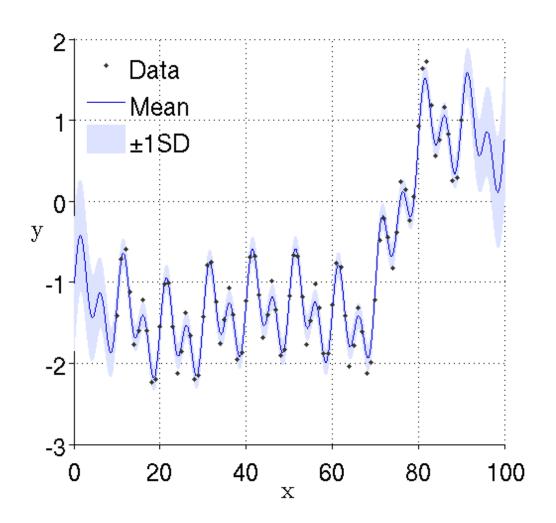
We can create new covariance functions by adding or multiplying other covariance functions.

e.g.





When a function is the sum of two independent functions, use a covariance that is the sum of the covariances for those two functions.





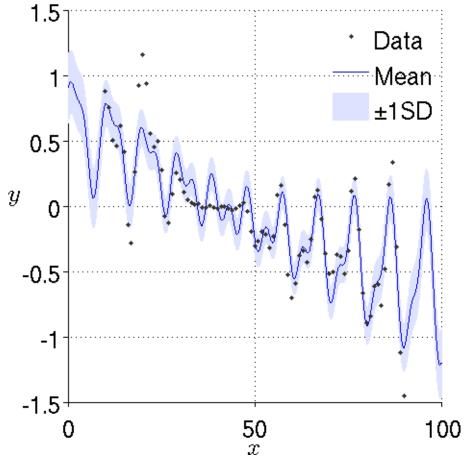
When a function is the product of two independent functions, use a covariance that is (almost) the product of the covariances for those two functions.

$$y(x) = a(x)b(x)$$

$$K_{y}(x_{1}, x_{2}) = K_{a}(x_{1}, x_{2})K_{b}(x_{1}, x_{2}) +$$

$$K_{a}(x_{1}, x_{2})\mu_{b}(x_{1})\mu_{b}(x_{2}) +$$

$$K_{b}(x_{1}, x_{2})\mu_{a}(x_{1})\mu_{a}(x_{2})$$



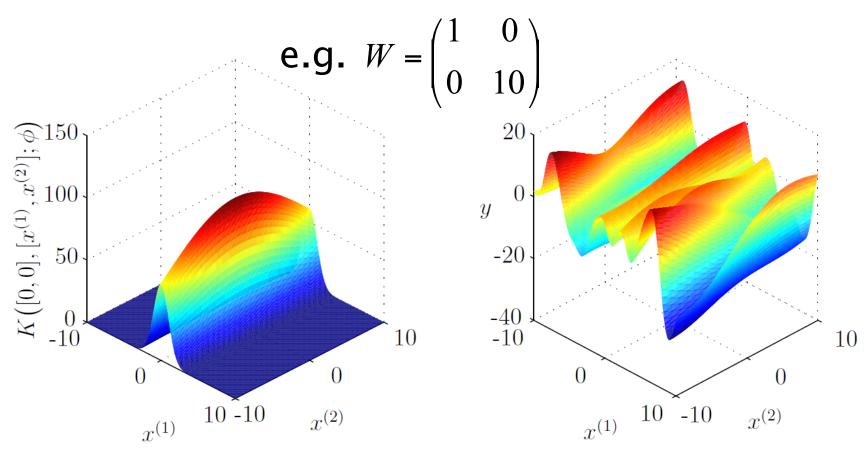


#### We can also modify covariance functions.

Inputs	Squared exponential	Matern	Polynomial
1-dim			(In (In)
n-dim			
derivative			
periodic  UNIVERSITY OF	TELL) TELL)		

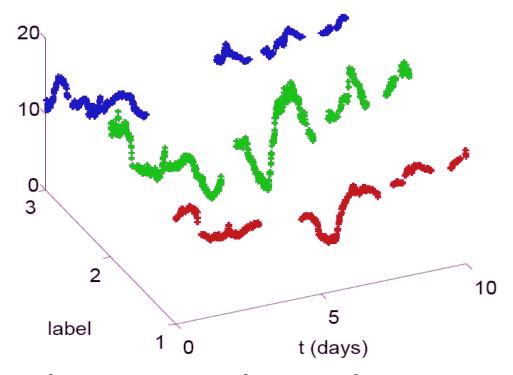
### We can modify covariance functions to accommodate multiple input dimensions, using

$$d(x_i, x_j; W) = \sqrt{(x_i - x_j)^T W^{-1}(x_i - x_j)}$$





If there are multiple outputs, reframe the problem as having a single output, and an additional *label* input specifying the output.



Hence we do not need simultaneous observations of all outputs.



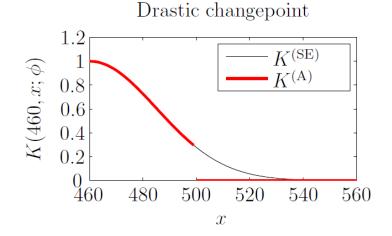
If the inputs were previously x, and outputs were labelled by l = 1, ..., L, we now need to specify a covariance over both x and l, e.g.

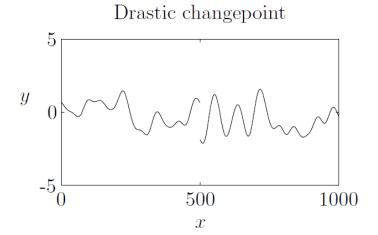
$$K((x_i, l_i), (x_j, l_j)) = K(x_i, x_j)K(l_i, l_j)$$

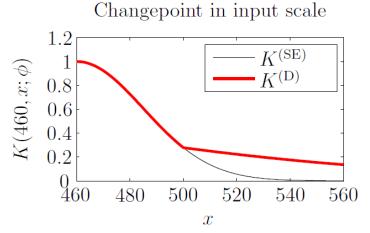
If *L* is not too large, we could use the spherical parameterisation.

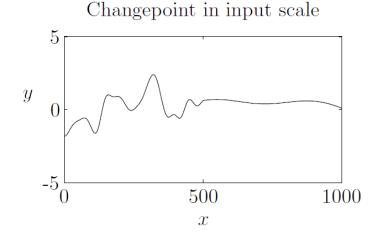


## Many other modifications are possible, to build covariances allowing for e.g. changepoints, faults and sets.









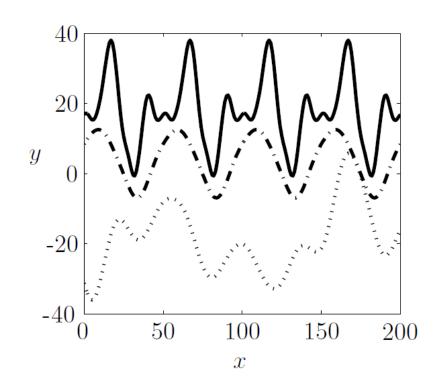


## We can modify covariance functions for functions known to be periodic, by using the distance

......Squared exponential Periodic sqd. exp., w = 1Periodic sqd. exp., w = 0.310080  $K(0, x; \phi)$ 60 40 20 0 50 100

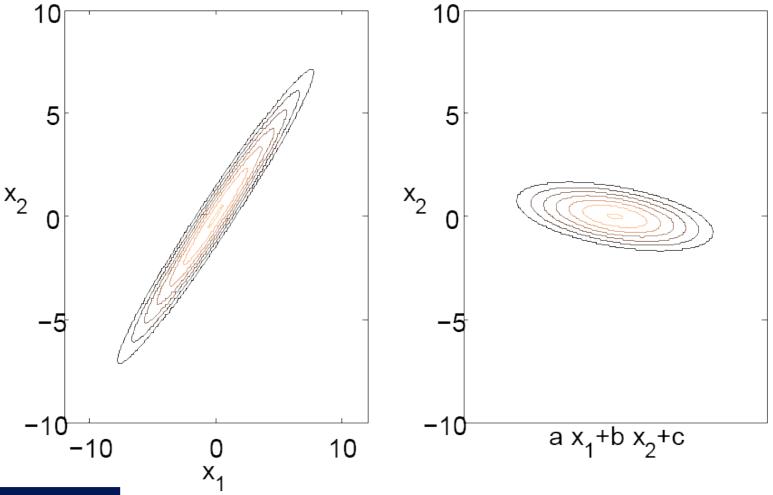
x

$$d(x_i, x_j) = \frac{1}{w} \left| \sin \left( \pi \frac{x_i - x_j}{T} \right) \right|$$



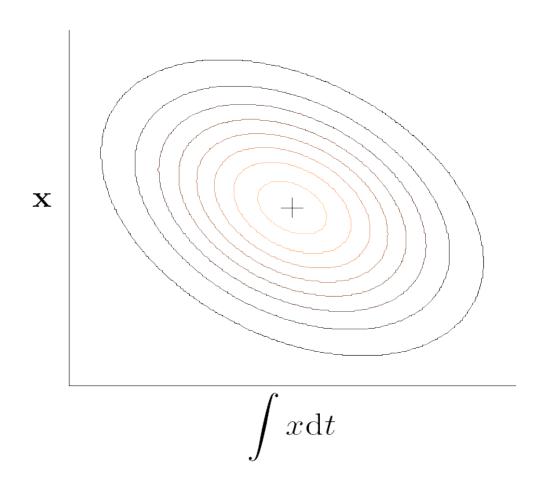


#### Gaussian distributed variables are joint Gaussian with any affine transform of them.





A function over which we have a Gaussian process is joint Gaussian with any integral or derivative of it, as integration and differentiation are affine.





### We can modify covariance functions to manage derivative or integral observations.

derivative observation at  $x_i$  and function observation at  $x_i$ .

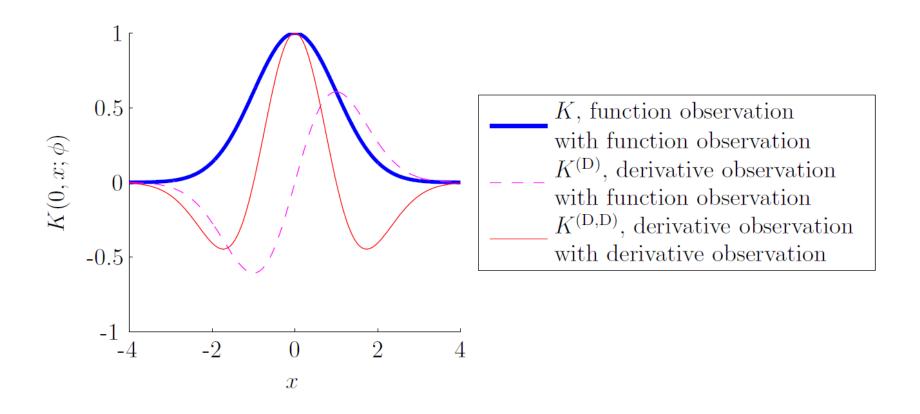
$$K_D(x_i, x_j) = \frac{\partial}{\partial x} K(x, x_j)\Big|_{x=x}$$

derivative observation at  $x_i$  and derivative observation at

$$K_{D,D}(x_i, x_j) = \frac{\partial}{\partial x'} \frac{\partial}{\partial x} K(x, x') \Big|_{x=x_i} \Big|_{x'=x_j}$$

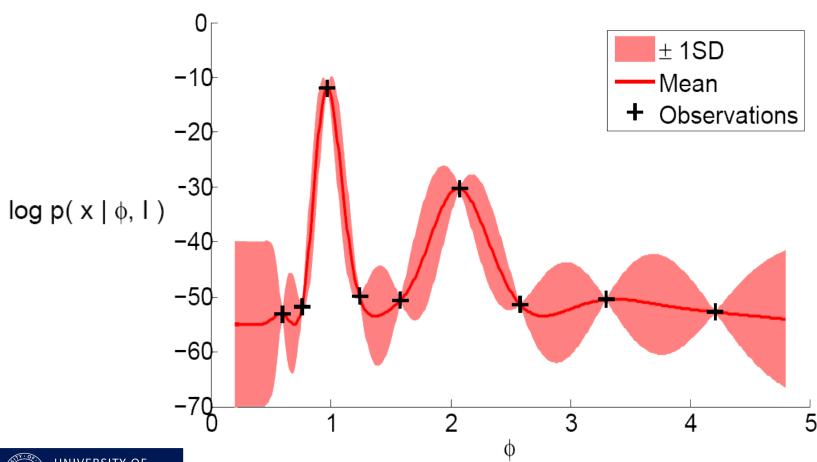


### We can modify the squared exponential covariance to manage derivative observations.



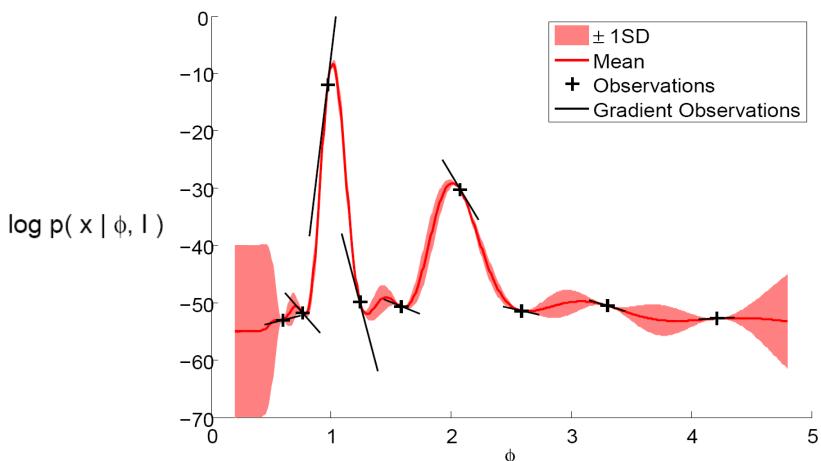


### We can improve our inference by including observations of the gradient of a function.



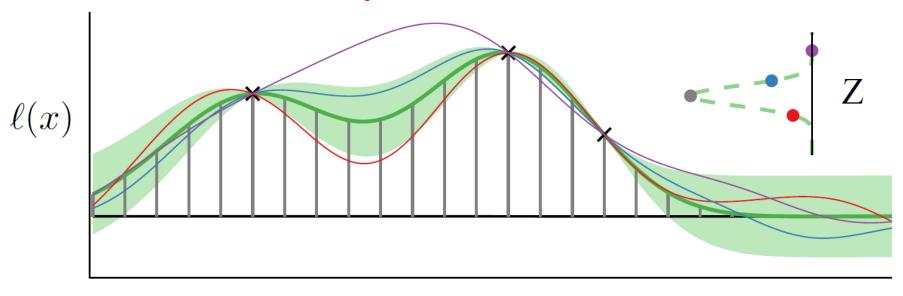


#### We can improve our inference by including observations of the gradient of a function.

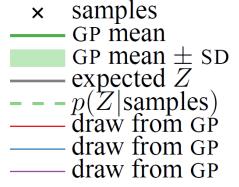




We can use observations of an integrand  $\ell$  in order to perform inference for its integral, Z: this is known as Bayesian Quadrature.



 $\mathcal{X}$ 

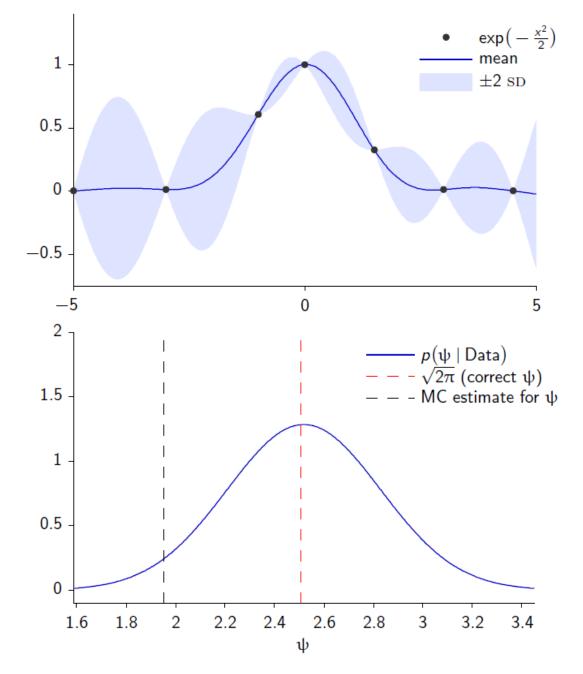




### Consider the integral

$$\psi = \int_{-5}^{5} \exp\left(-\frac{x^2}{2}\right) \mathrm{d}x$$

Bayesian quadrature achieves more accurate results than Monte Carlo, and provides an estimate of our uncertainty.





The enormous flexibility afforded by covariance functions comes at a price: hyperparameters, which must be marginalised.

$$p(y_{\star}|\mathbf{z}_d) = \frac{\int p(y_{\star}|\mathbf{z}_d, \phi) p(\mathbf{z}_d|\phi) p(\phi) d\phi}{\int p(\mathbf{z}_d|\phi) p(\phi) d\phi}$$

Unfortunately, these integrals are non-analytic.

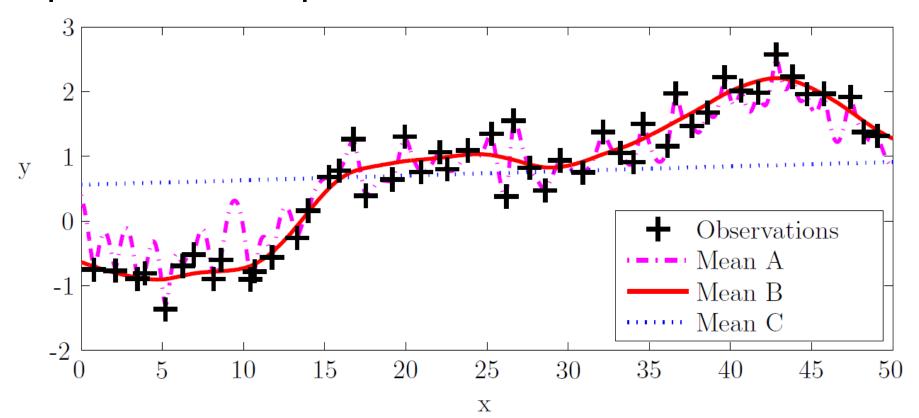


Given that we don't to fix  $y_*$ , the two important terms in our integrands are the likelihood and the prior (specifically, their product, proportional to the posterior for  $\phi$ ).

$$p(y_{\star}|\boldsymbol{z}_{d}) = \frac{\int p(y_{\star}|\boldsymbol{z}_{d},\phi)p(\boldsymbol{z}_{d}|\phi)p(\phi) d\phi}{\int p(\boldsymbol{z}_{d}|\phi)p(\phi) d\phi}$$



The hyperparameter priors can have a significant influence on our inference. Prior A favours small input scale, prior C favours large input scale and prior B is uninformative.





Selecting priors is easy.

#### **Prior**

Use what you know.



If probability theory makes 'wrong' predictions, then we have learned something!

Model (/)



Probability theory

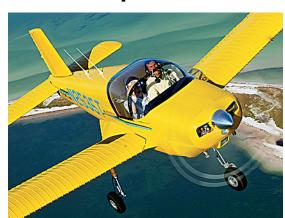


Our expectations

**Predictions** 









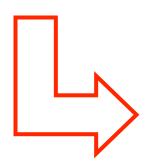
If probability theory makes 'wrong' predictions, then we have learned something!

One of these Model (1) two must be wrong **Probability** theory Our expectations **Predictions #** 



In this way, we are led to construct successively better models.

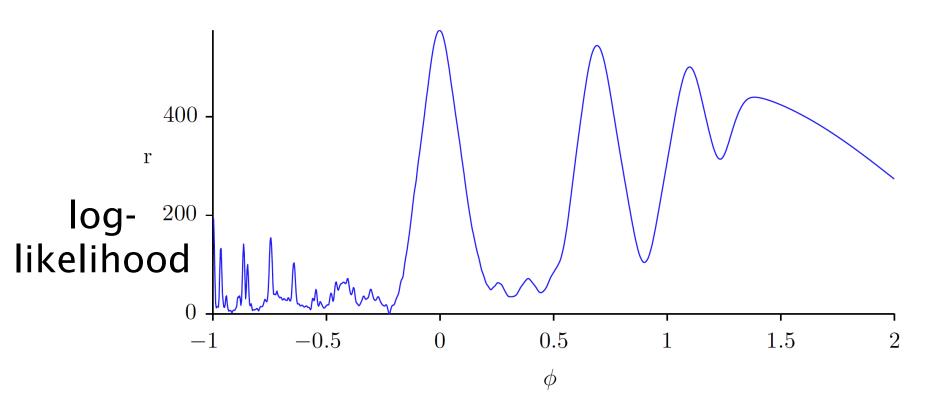








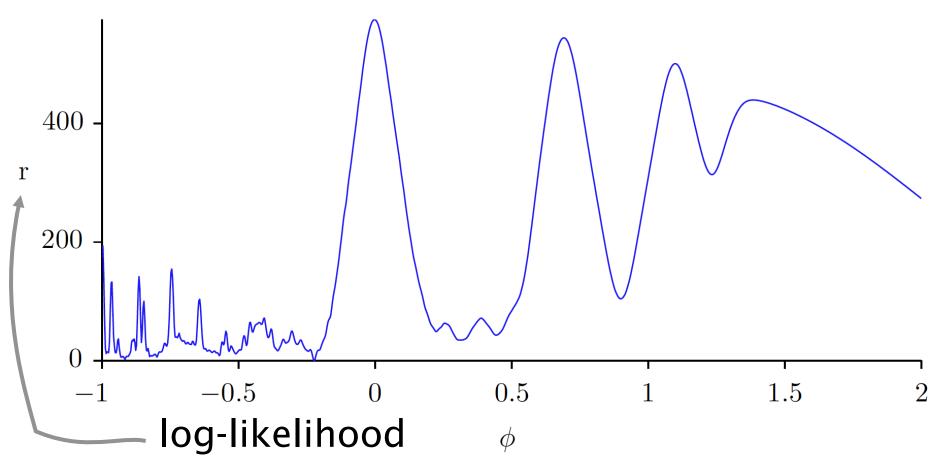
Marginalisation requires quadrature, which presents two challenges: integrand exploration, and integral estimation.



hyperparameters

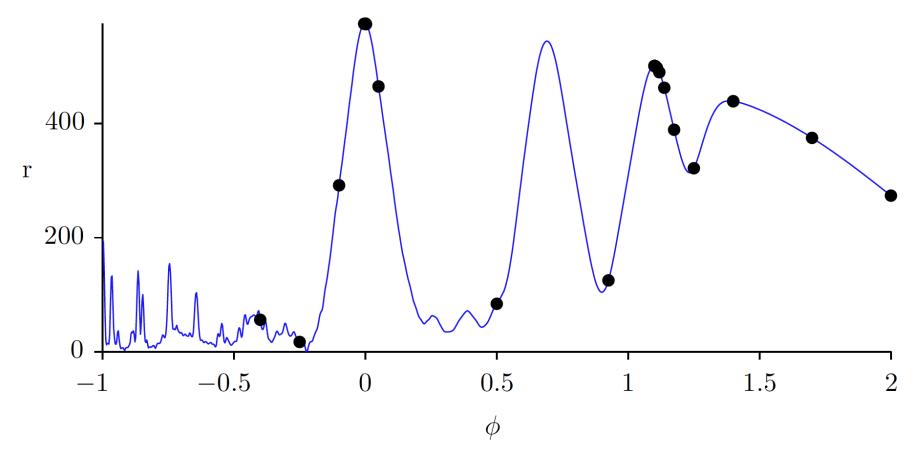


There are many different approaches to quadrature for probabilistic integrals; integrand estimation is usually undervalued.



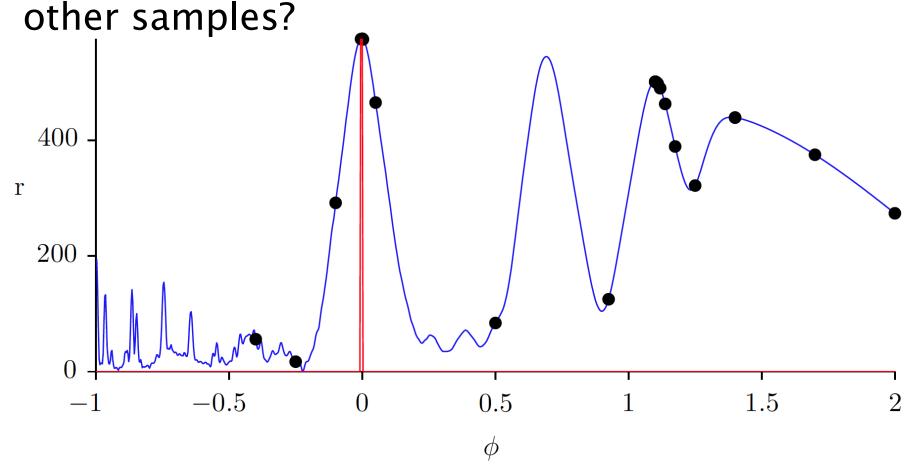


Optimisation (as in maximum likelihood), particularly using global optimisers, gives a reasonable heuristic for exploring the integrand.



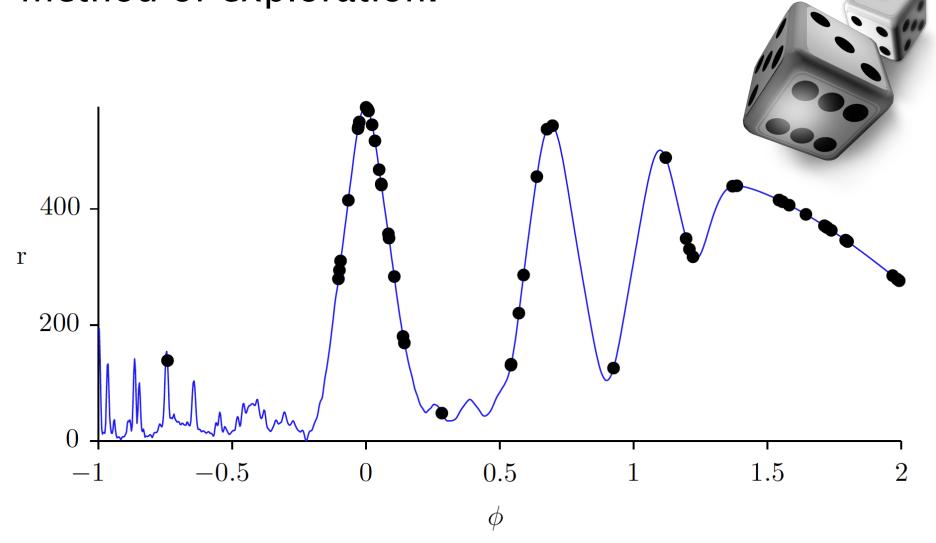


However, maximum likelihood is an unreasonable way of estimating a multi-modal likelihood integrand: why throw away all those other camples?

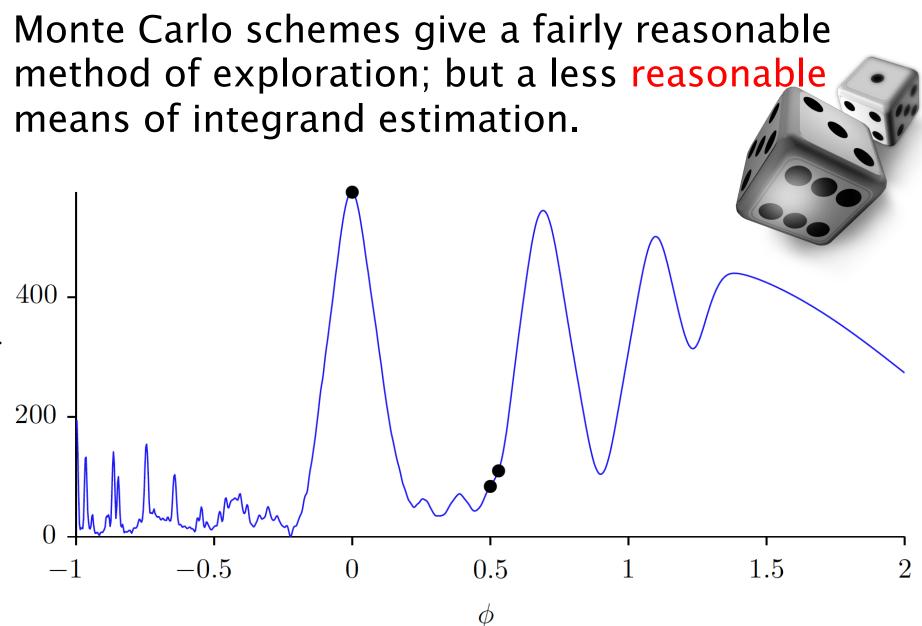




Monte Carlo schemes give a another reasonable method of exploration.

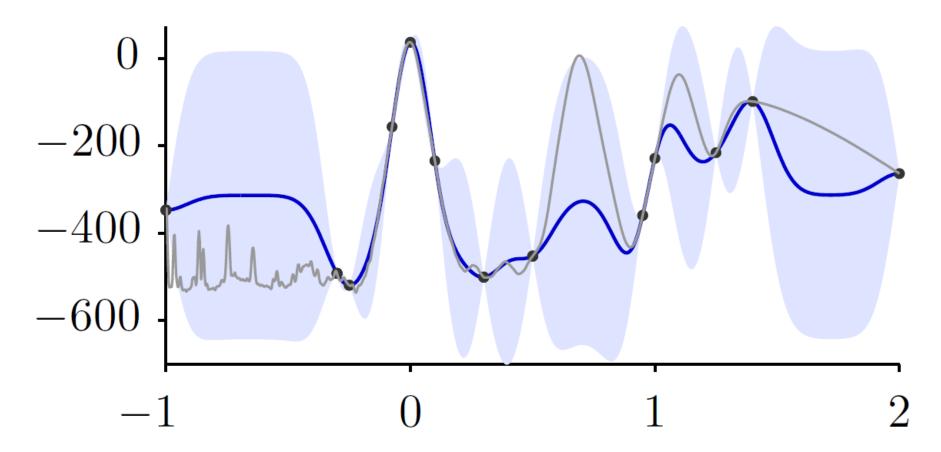






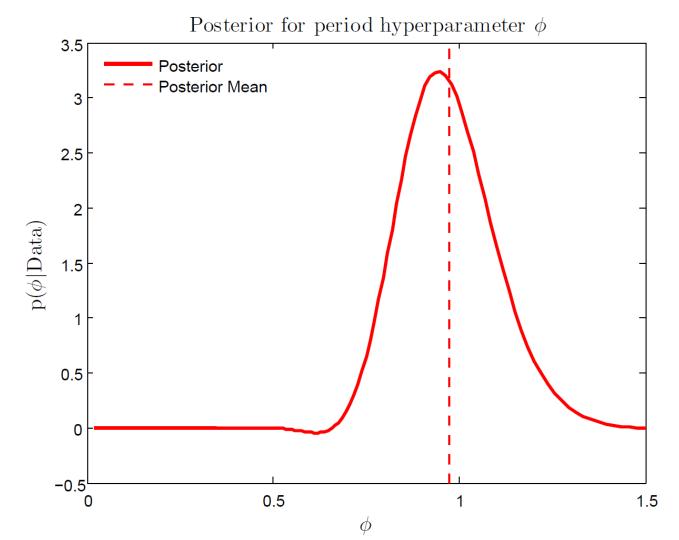


Bayesian Monte Carlo uses samples obtained via Monte Carlo within a Bayesian quadrature framework to give an estimate for the integral.





#### With Bayesian quadrature, we can also estimate the posterior distributions for any hyperparameters.



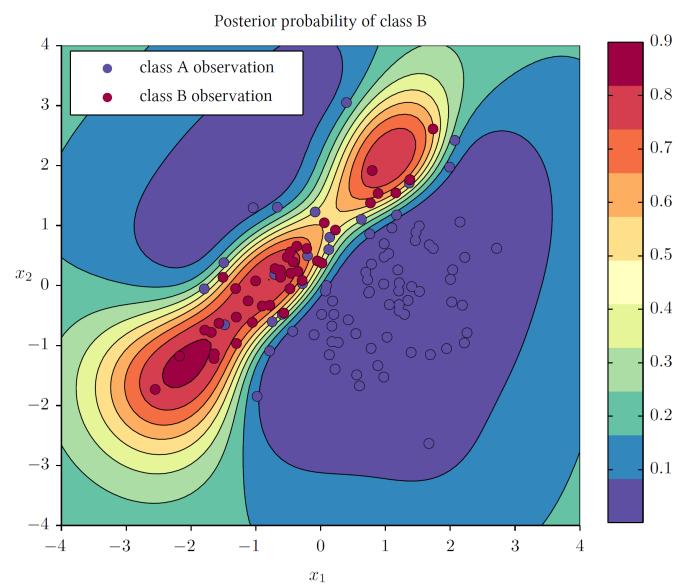


There are many approaches to hyperparameter marginalisation, but only two are recommended.

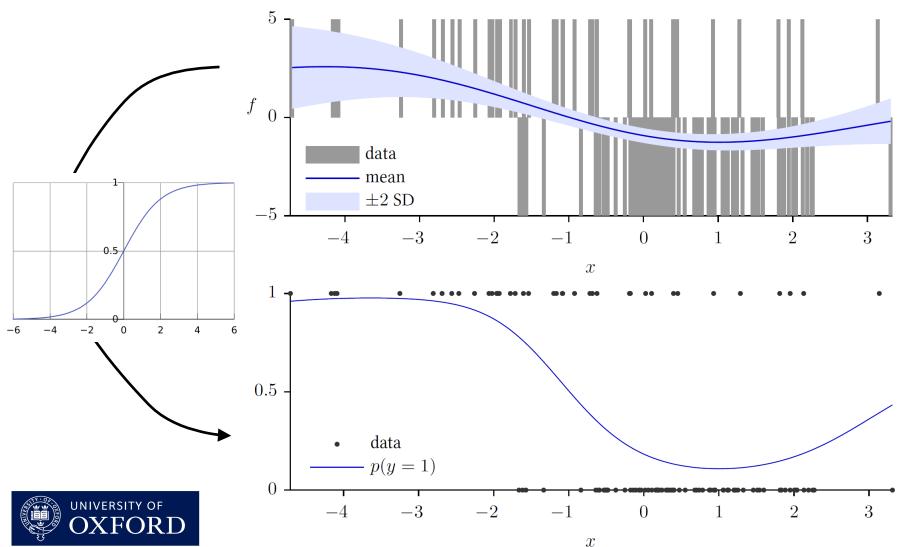
Likelihood	Marginalisation
Unimodal or high dimensional	Maximum likelihood.
Multimodal or computationally expensive	Bayesian Monte Carlo.



## We can put Gaussian processes to work not just for regression. but also for classification.



To do so, use a Gaussian process to model a latent variable, mapped through a sigmoid to a discrete class label.

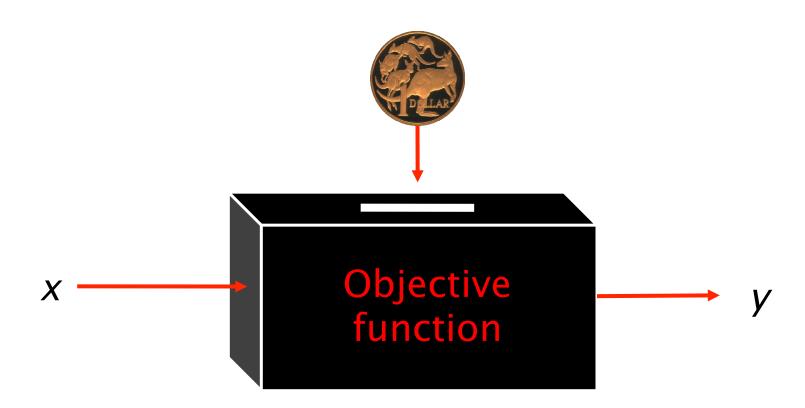


Unfortunately, using this sigmoid makes inference intractable. Approximate inference can be achieved using a number of algorithms.

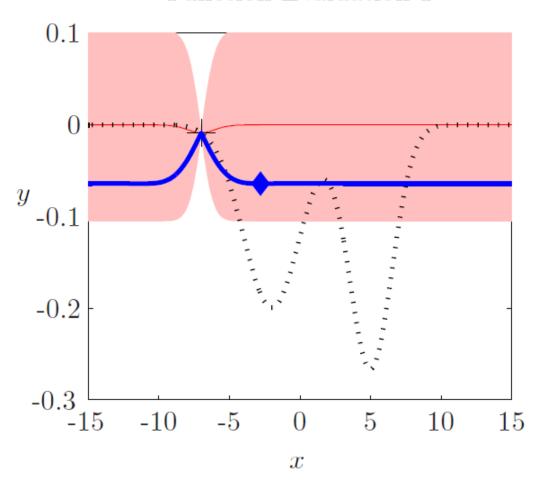
Algorithm	Speed	Accuracy
Laplace approximation	Very fast	Low.
Expectation Propagation	Fast	High.
Markov Chain Monte Carlo	Very slow	Very high.

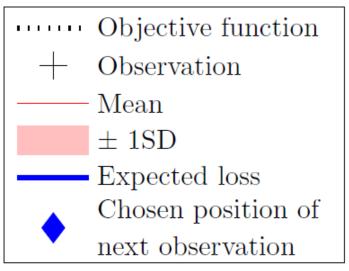


We treat global optimisation as a Bayesian decision problem.

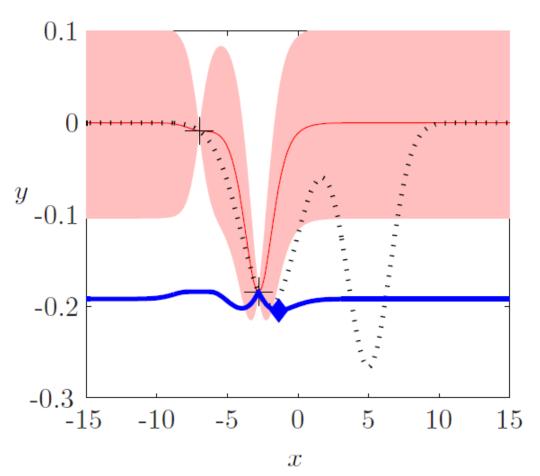


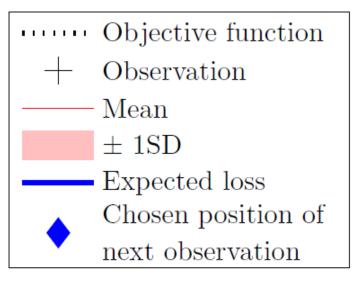




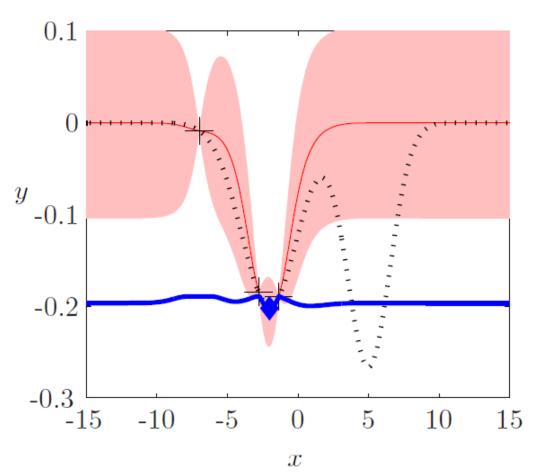


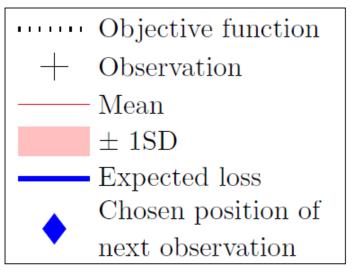




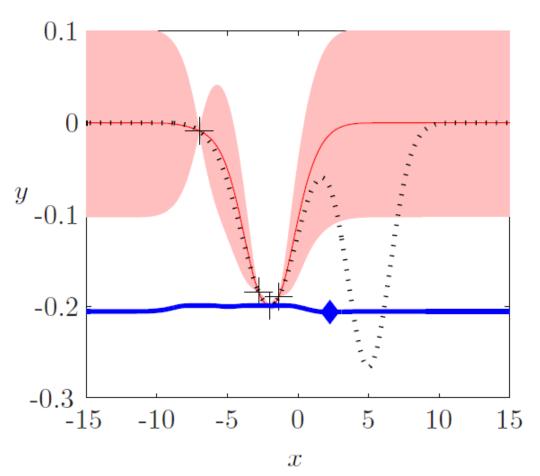


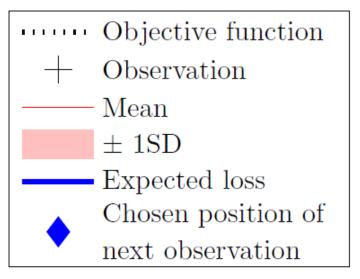




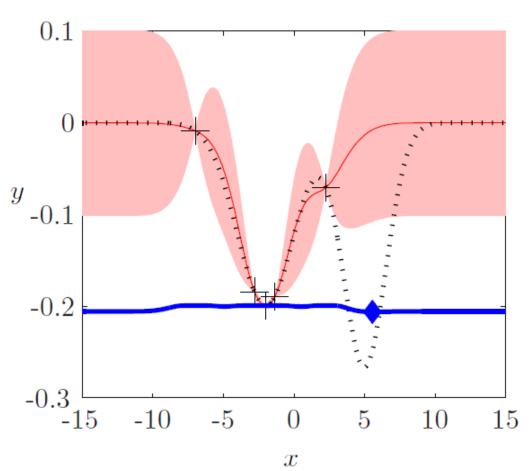


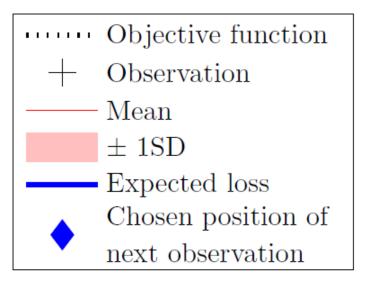




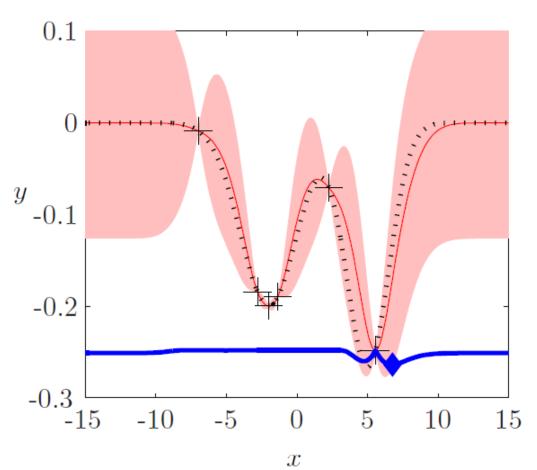


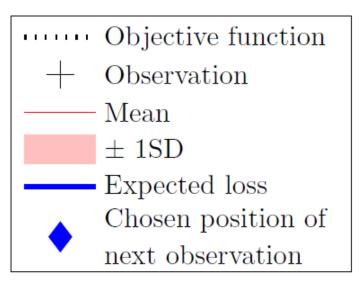




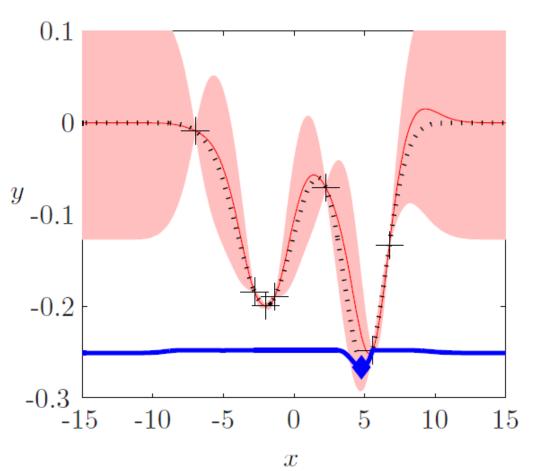


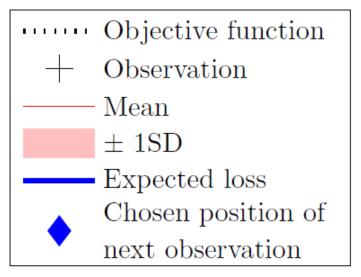




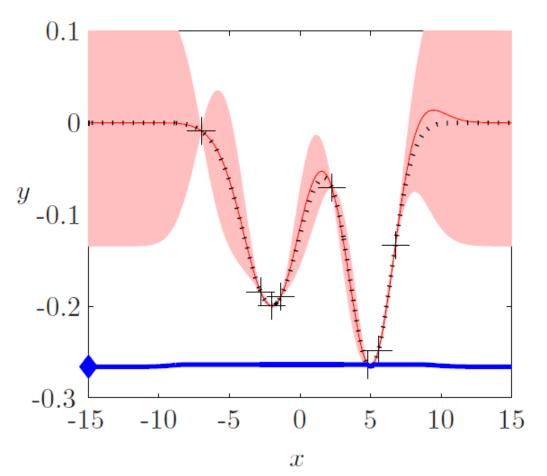


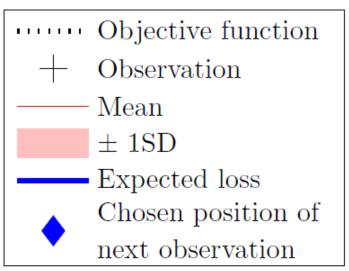




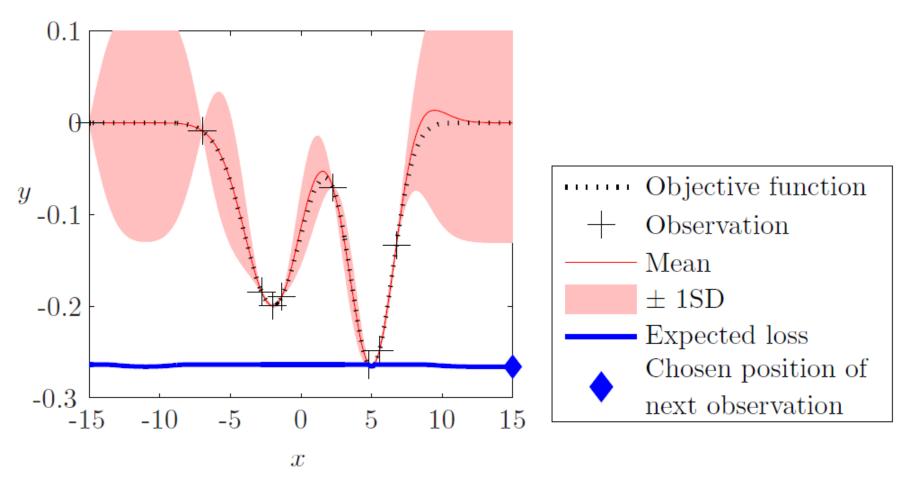














The key computational bottleneck associated with Gaussian processes is resolving inv(K) v, or, equivalently, solving v = K x for x.

$$egin{pmatrix} K_{11} & K_{21} & K_{13} & \cdots \\ K_{21} & K_{22} & K_{23} & & & v_2 \\ K_{31} & K_{32} & K_{33} & & & v_3 \\ dots & & \ddots \end{pmatrix}^{-1} egin{pmatrix} v_1 \\ v_2 \\ v_3 \\ dots \\ dots \end{pmatrix}$$

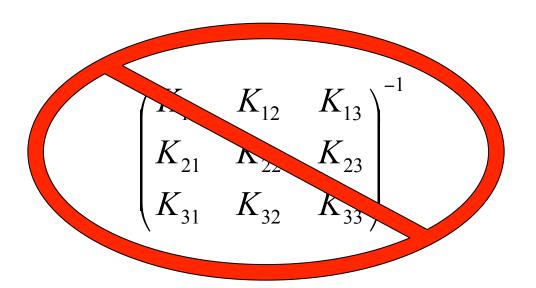


Our choice of a method to solve v = K x for x depends on the structure of covariance K.

<b>Covariance matrix</b>	Solving method
Poorly conditioned	Improve conditioning, then see below.
(Just) positive semi- definite	Cholesky factorisation.
Toeplitz	Toeplitz solver.
Kronecker product	Kronecker solver.
Too big and dense	Sparsification.
Updated version of previous matrix	Update, dependent on above.



You should never actually invert a matrix.



Inversion is slow,  $O(n^3)$  in matrix size n.

Inversion is also unstable; conditioning errors are significant.



Conditioning becomes an issue when we have multiple close observations, giving rows in the covariance matrix that are very similar.

The condition number (cond) of a covariance matrix is the ratio of the largest to the smallest eigenvalue; in Matlab, things break down at about cond(K) =  $10^{16}$ .

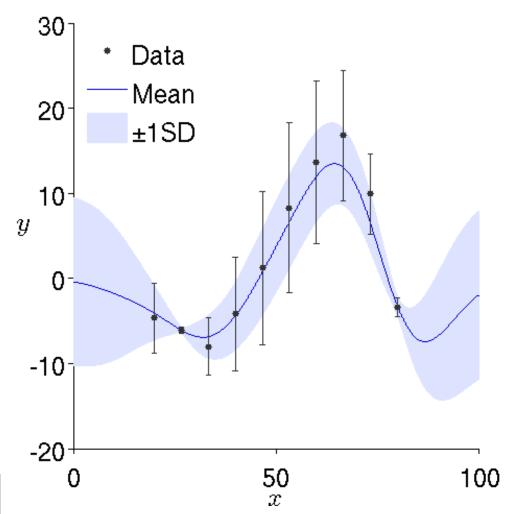


The solution to conditioning problems is to add a small positive quantity (*jitter*) to the diagonal of the covariance matrix.

1.01	0.9999	0	0	☐ Sufficiently dissimilar
0.9999	1.01	0	0	
0	0	1.01	0.1	
0	0	0.1	1.01	



As jitter is effectively imposed noise, adding jitter to all diagonal elements (unnecessarily) dilutes the informativeness of our data.





The Cholesky factorisation of a positive semidefinite matrix K is relatively fast  $(1/3 \text{ O}(n^3) \text{ in }$ matrix size n) and more numerically stable.

$$K = R^{\mathrm{T}}R$$

$$R = \text{chol}(K) = \begin{pmatrix} R_{11} & R_{12} & \cdots & R_{1n} \\ 0 & R_{22} & \cdots & R_{2n} \\ \vdots & \vdots & \ddots & \vdots \\ 0 & 0 & \cdots & R_{nn} \end{pmatrix}$$



The upper triangular Cholesky factor can then be stored and used to solve v = K x for x very quickly (O( $n^2$ ) in matrix size n) by back substitution.

$$v = Kx$$

$$v = R^{T}x'$$

$$x' = Rx$$

$$\begin{pmatrix} v_{1} \\ v_{2} \\ \vdots \\ v_{n} \end{pmatrix} = \begin{pmatrix} R_{11} & R_{12} & \cdots & R_{1n} \\ 0 & R_{22} & \cdots & R_{2n} \\ \vdots & \vdots & \ddots & \vdots \\ 0 & 0 & \cdots & R_{nn} \end{pmatrix} \begin{pmatrix} x'_{1} \\ x'_{2} \\ \vdots \\ x'_{n} \end{pmatrix}$$



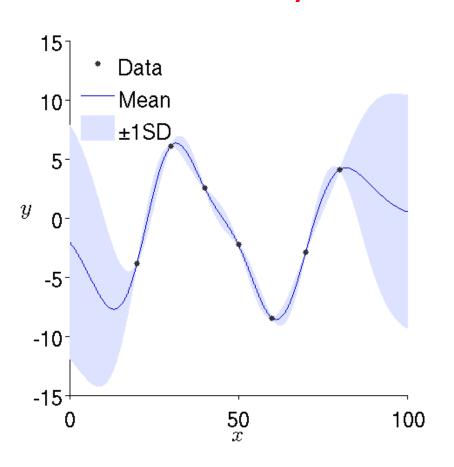
A symmetric matrix *K* is Toeplitz if it can be written as

$$K = \begin{pmatrix} k_1 & k_2 & k_3 & k_4 & \cdots & k_n \\ k_2 & k_1 & k_2 & k_3 & & & \\ k_3 & k_2 & k_1 & k_2 & & & \\ k_4 & k_3 & k_2 & k_1 & & & \\ \vdots & & & \ddots & & \\ k_n & & & & k_1 \end{pmatrix}$$

If K is Toeplitz, there exists a very efficient method to solve v = K x for x (O(4 $n^2$ ) in matrix size n).



# A Gaussian process has a Toeplitz covariance matrix if we have linearly spaced observations and a stationary covariance function.



$$K = \begin{pmatrix} 100 & 60.7 & 13.5 & 1.11 & 0.03 & 0 & 0 \\ 60.7 & 100 & 60.7 & 13.5 & 1.11 & 0.03 & 0 \\ 13.5 & 60.7 & 100 & 60.7 & 13.5 & 1.11 & 0.03 \\ 1.11 & 13.5 & 60.7 & 100 & 60.7 & 13.5 & 1.11 \\ 0.03 & 1.11 & 13.5 & 60.7 & 100 & 60.7 & 13.5 \\ 0 & 0.03 & 1.11 & 13.5 & 60.7 & 100 & 60.7 \\ 0 & 0 & 0.03 & 1.11 & 13.5 & 60.7 & 100 \end{pmatrix}$$



## Some special large matrices can be represented in a compact way using the Kronecker product.

$$\mathbf{A} \otimes \mathbf{B} = \begin{bmatrix} a_{11}B & \cdots & a_{1n}B \\ \vdots & \ddots & \vdots \\ a_{m1}B & \cdots & a_{mn}B \end{bmatrix}.$$

$$\mathbf{A} \otimes \mathbf{B} = \begin{bmatrix} a_{11}b_{11} & a_{11}b_{12} & \cdots & a_{11}b_{1q} & \cdots & \cdots & a_{1n}b_{11} & a_{1n}b_{12} & \cdots & a_{1n}b_{1q} \\ a_{11}b_{21} & a_{11}b_{22} & \cdots & a_{11}b_{2q} & \cdots & \cdots & a_{1n}b_{21} & a_{1n}b_{22} & \cdots & a_{1n}b_{2q} \\ \vdots & \vdots & \ddots & \vdots & & \vdots & \ddots & \vdots \\ a_{11}b_{p1} & a_{11}b_{p2} & \cdots & a_{11}b_{pq} & \cdots & \cdots & a_{1n}b_{p1} & a_{1n}b_{p2} & \cdots & a_{1n}b_{pq} \\ \vdots & \vdots & & \vdots & \ddots & & \vdots & & \vdots \\ \vdots & \vdots & & \vdots & & \ddots & & \vdots & & \vdots \\ a_{m1}b_{11} & a_{m1}b_{12} & \cdots & a_{m1}b_{1q} & \cdots & \cdots & a_{mn}b_{11} & a_{mn}b_{12} & \cdots & a_{mn}b_{1q} \\ a_{m1}b_{21} & a_{m1}b_{22} & \cdots & a_{m1}b_{2q} & \cdots & \cdots & a_{mn}b_{21} & a_{mn}b_{22} & \cdots & a_{mn}b_{2q} \\ \vdots & \vdots & \ddots & \vdots & & \vdots & \ddots & \vdots \\ a_{m1}b_{p1} & a_{m1}b_{p2} & \cdots & a_{m1}b_{pq} & \cdots & \cdots & a_{mn}b_{p1} & a_{mn}b_{p2} & \cdots & a_{mn}b_{pq} \end{bmatrix}$$



If K is a Kronecker product, there exists a very efficient method to solve v = K x for x (particularly when v is itself a Kronecker product):

size 
$$n_a \times n_b$$
  

$$x = (K_a \otimes K_b)^{-1}(v_a \otimes v_b)$$

$$= (K_a^{-1}v_a) \otimes (K_b^{-1}v_b)$$

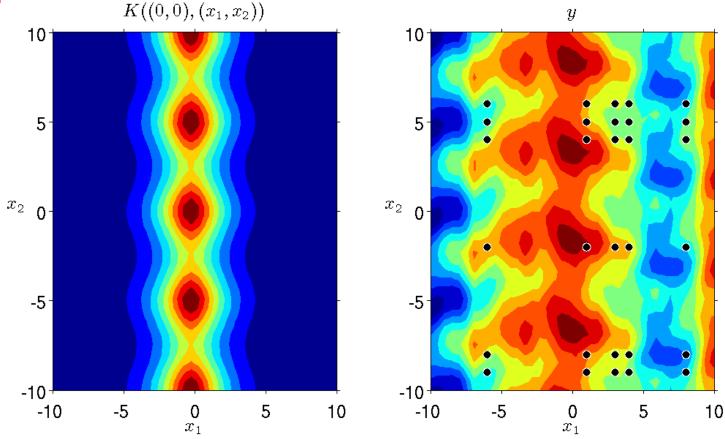
$$= (K_a \otimes n_a \otimes k_b) \otimes (K_b \otimes k_b)$$
size  $n_a \otimes k_b$ 

Recall that solving operations are typically O  $(n^3)!$ 



A Gaussian process will have a Kronecker product for a covariance matrix if we use a product covariance function and a grid of

samples.





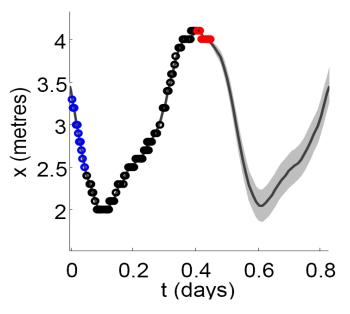
If a very large covariance matrix is not decomposable as a Kronecker product (or otherwise), we may wish to attempt sparsification.

$$K = \begin{pmatrix} K_{11} & K_{12} & 0 & 0 & \cdots & 0 \\ K_{21} & K_{22} & K_{23} & 0 & & & \\ 0 & K_{32} & K_{33} & K_{34} & & & \\ 0 & 0 & K_{43} & K_{44} & & & \\ \vdots & & & \ddots & & \\ 0 & & & & K_{nn} \end{pmatrix}$$



There are many ways to sparsify our data; the simplest involve selecting a subset. Windowing represents a reasonable way to do this.

Tide Heights, Independent Sensor 1



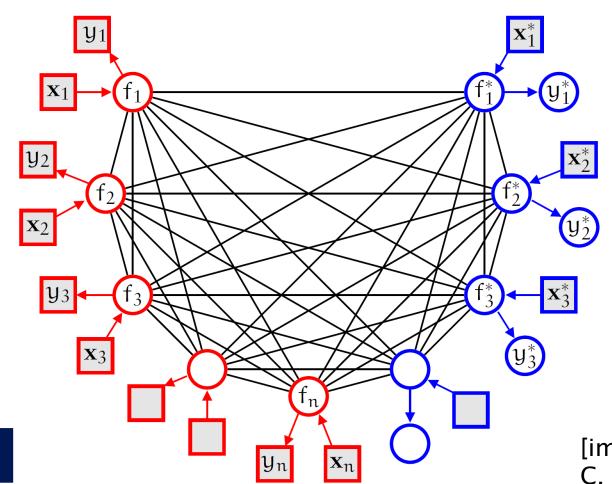
± 1SD

Mean

- Discarded Observations
- Old Observations
- New Observations



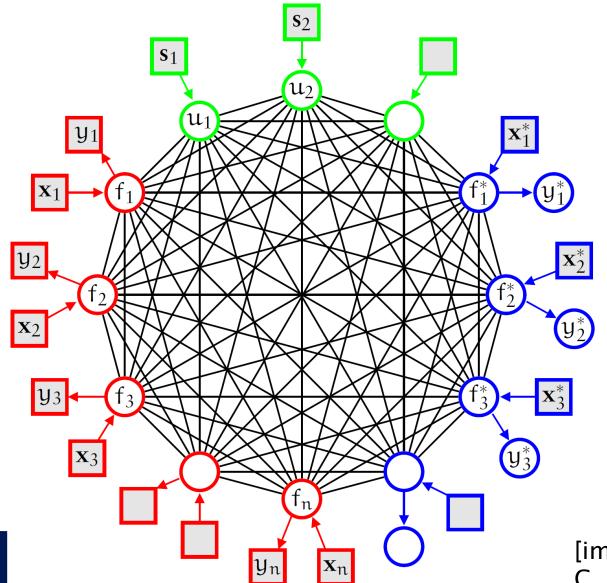
#### A Gaussian process assumes all variables f are correlated.





[image courtesy C. Rasmussen]

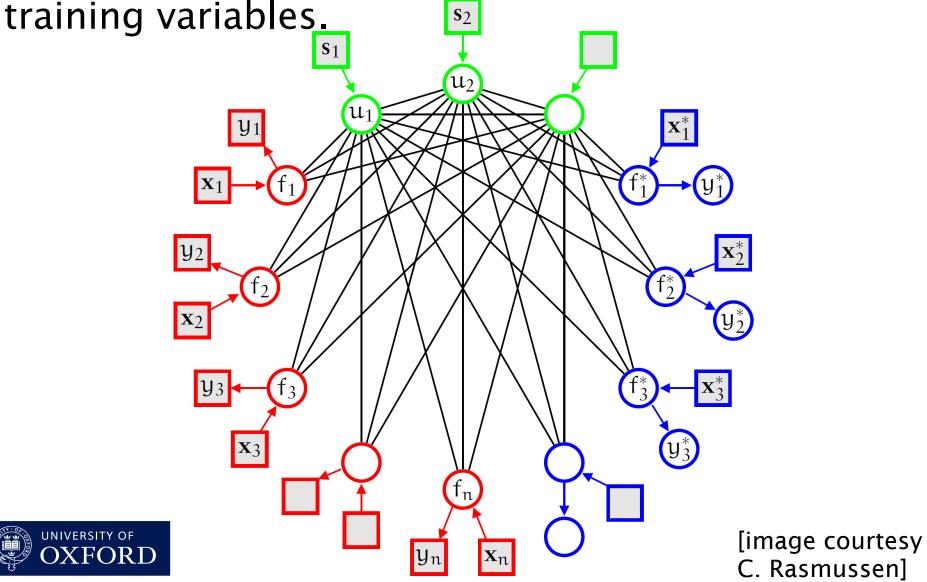
Imagine introducing additional, unobserved inducing variables *u*.



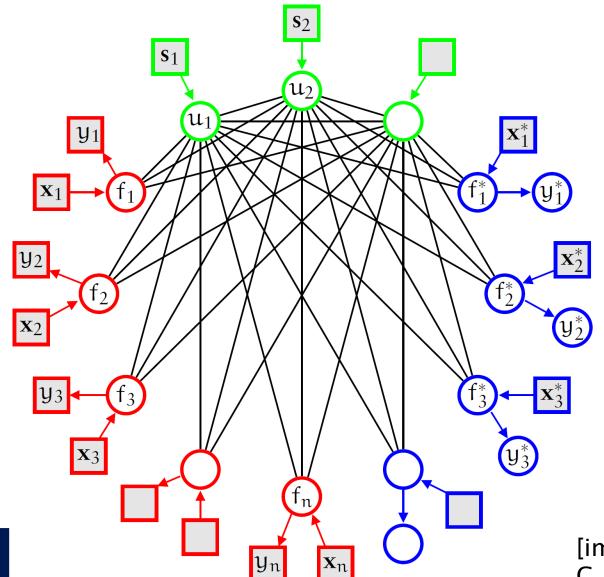


[image courtesy C. Rasmussen]

We can sparsify data by using inducing variables to mediate the interactions between test and



There are many such schemes for sparsification, that differ in the choice of inducing inputs.





[image courtesy C. Rasmussen]

Finally, if we already have the Cholesky factor

$$R_{11} = \operatorname{chol}(K_{11}),$$

we can efficiently determine the updated factor

$$\begin{pmatrix} R_{11} & R_{12} \\ 0 & R_{22} \end{pmatrix} = \text{chol} \begin{pmatrix} K_{11} & K_{12} \\ K_{12} & K_{22} \end{pmatrix},$$

and similar for other types of Cholesky updates and downdates, and for solutions based upon them. A Toeplitz update is probably also possible.



We want to evaluate a large number of hyperparameter samples to explore hyperparameter space. Fortunately, each sample can be evaluated in parallel (possibly on a graphics card).





I hope you have learned how to fit Gaussian processes to data.

